



# Overview & Perspectives "Ab Initio & EFT Programs"

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# Projects

- **A02: Effective field theories and ab initio calculations of light nuclei**

PIs: Hans-Werner Hammer, Robert Roth

- **A04: Strong interactions and structure of medium-mass nuclei**

PIs: Robert Roth, Achim Schwenk

- **B05: Nuclear matter equation of state for astrophysical applications**

PIs: Jens Braun, Kai Heleler

- **A05: Halos and clustering in nuclei**

PIs: Tom Aumann, Pierre Capel, Hans-Werner Hammer

- **B01: Electroweak interactions in nuclei and nuclear matter**

PIs: Gabriel Martínez-Pinedo, Achim Schwenk

- **B04: Electric dipole response and neutron equation of state**

PIs: Sonia Bacca, Heiko Scheit

# Overview

**A02**

**A04**

**B05**

**Domain**

light nuclei (s & p-shell)  
 ${}^3n, {}^4n$  systems

medium-mass &  
heavy nuclei  
open-shell systems

infinite nuclear/neutron  
matter

**Frontiers**

precision calculations of  
elmag. observables  
inclusion of continuum  
resonance physics

observables beyond  
ground-states  
multi-reference/deformation  
heavy nuclei

extension & merging to  
high-density regime  
finite temperature  
arbitrary proton fraction

**Uncertainty Quantification**

**Many-Body Tools**

**(Chiral) Effective Field Theory**

# Overview

A02

A04

B05

**A05**

**B01**

**B04**

halo nuclei  
nn system

electroweak transitions  
neutrino-matter interactions

electric dipole polarizability

emag. observables  
inclusion of continuum  
resonance physics

ground-state energies  
multi-reference/deformation  
heavy nuclei

high-density regime  
finite temperature  
arbitrary asymmetry

**Uncertainty Quantification**

**Many-Body Tools**

**(Chiral) Effective Field Theory**

Domain

Frontiers

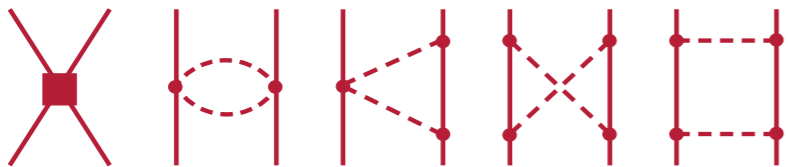
# Nuclear Interactions from Chiral EFT

## NN

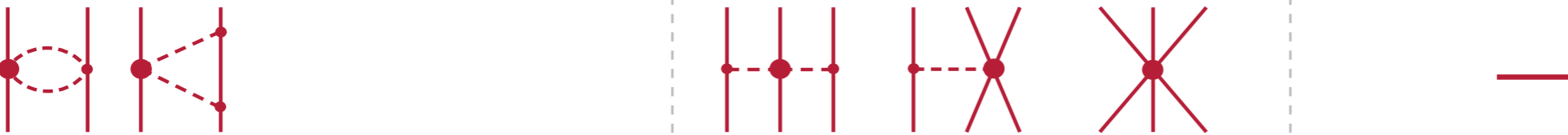
LO



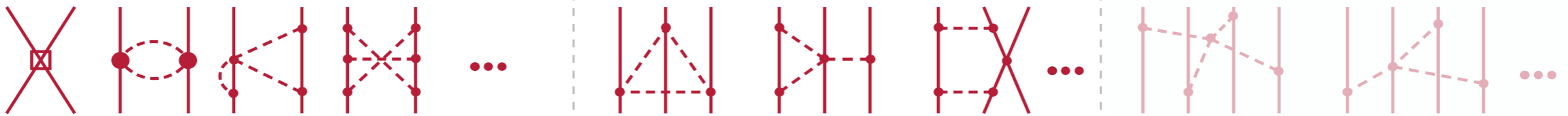
NLO



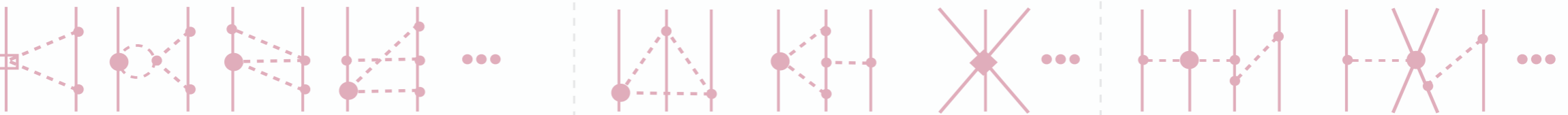
N2LO



N3LO



N4LO



## Non-Local Interactions

- NN+3N interactions with same chiral orders and regulator scheme and scale
- up to N3LO with different cutoff values
- low-energy constants fit to  $A=2,3$  data and  $^{16}\text{O}$  ground-state energy

*T. Hüther et al., PLB 808, 135651 (2020)*

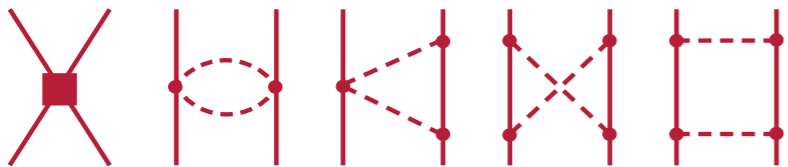
# Nuclear Interactions from Chiral EFT

NN

LO



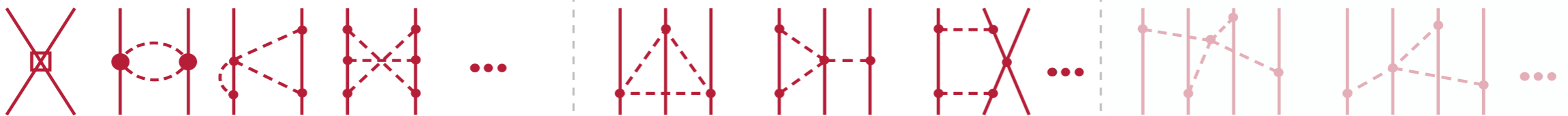
NLO



N2LO



N3LO



N4LO



## Semi-Local Interactions

- consistent NN+3N interactions and two-body currents / charge densities
- NN available up to N4LO+  
3N at N2LO... soon N3LO, partial N4LO
- low-energy constants fit exclusively to  $A=2,3$  data

*LENPIC Collab, PRC 103, 054001 (2021)  
PRC (2022) accepted*

# Uncertainty Quantification

all theory calculations are affected by systematic uncertainties (some also by statistical uncertainties)

in ab initio methods uncertainties result in a controlled way from truncations

variation of all truncation parameters gives access to systematic UQ

## ■ **chiral EFT uncertainties**

- assess convergence of observables in dependence of chiral order
- convergence affected by regulator scheme and scale

use Bayesian methods to estimate size of omitted contributions based on known systematics

## ■ **many-body uncertainties**

- assess convergence of observables with model-space truncation
- convergence affected by basis choice, truncation scheme, Hamiltonian

# Many-Body Tools

**A02**

**A04**

**B05**

Few-Body Methods

Model / Configuration Interaction

Shell Model

Perturbation Theory

Similarity Renormalization

In-Medium Similarity Renormalization Group

Functional RG

**all methods are computationally limited...**

...



# Many-Body Tools: Emulators

use emulators to circumvent computationally expensive many-body calculations

real many-body calculations provide the information needed for the construction of the emulator

## ■ **eigenvector continuation for parameter variations**

- use tiny but highly problem-adapted basis set extracted from full many-body solution (eigenvectors)
- powerful tool for large-scale parameter variations, e.g. for LEC fitting

## ■ **artificial neural networks for model-space extrapolation**

- train ANN to predict converged observables based on sequences of non-converged many-body calculations
- provides robust "extrapolations" with statistically sound uncertainties

■ ■ ■

Just one specific example...

# A04: In-Medium NCSM Study of Neon Isotopes

# No-Core Shell Model

*Barrett, Vary, Navrátil, Maris, Nogga, Roth,...*

no-core shell model is  
universal and powerful ab initio approach for  
light nuclei (up to  $A \approx 25$ )

- **idea**: solve eigenvalue problem of Hamiltonian represented in model space of HO Slater determinants truncated w.r.t. HO excitation energy  $N_{\max} \hbar \Omega$

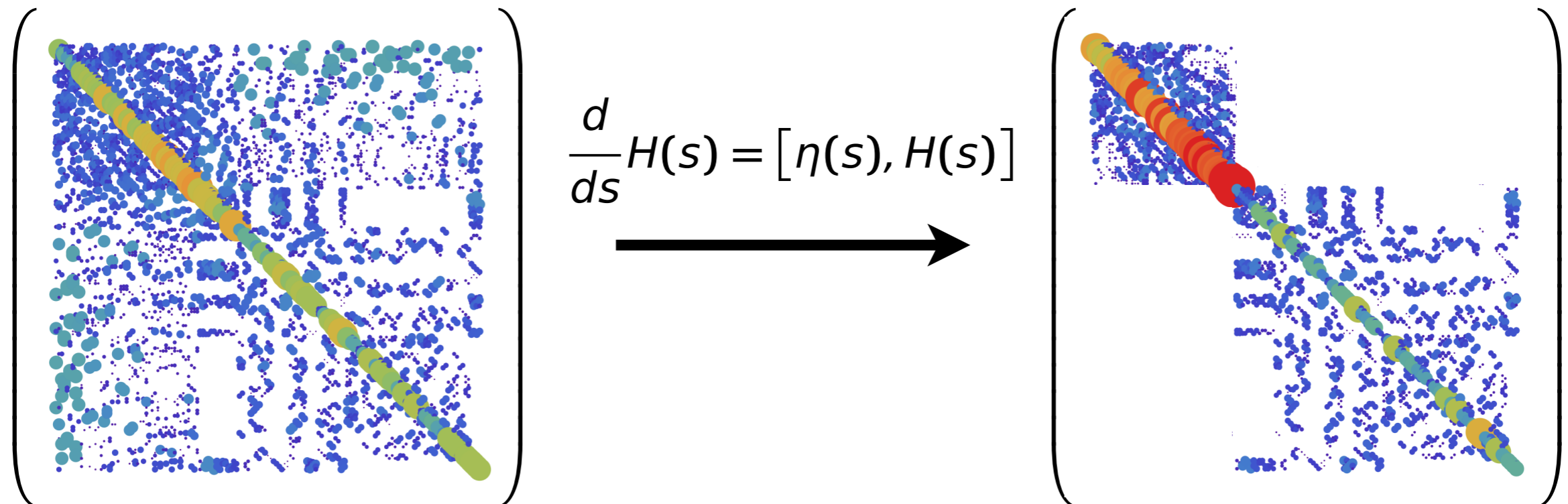
$$\left( \begin{array}{c} \text{[Matrix of blue dots with a diagonal band of yellow and orange dots]} \end{array} \right) \begin{pmatrix} \vdots \\ C_{i'}^{(n)} \\ \vdots \end{pmatrix} = E_n \begin{pmatrix} \vdots \\ C_i^{(n)} \\ \vdots \end{pmatrix}$$

# Multi-Reference In-Medium SRG

Hergert, Gebrerufael, Vobig, Mongelli, Roth,...

decouple reference state from excitations by a unitary transformation of Hamiltonian and other operators

- **idea**: use multi-reference formulation of IM-SRG to decouple reference space for rest of model space, i.e., block diagonalize  $A$ -body Hamiltonian



# In-Medium NCSM

**NCSM**  
reference state

- ground-state from NCSM at small  $N_{\max}$  as reference state for multi-reference IM-SRG
- access to all open-shell nuclei and systematically improvable

**MR-IM-SRG**  
decoupling

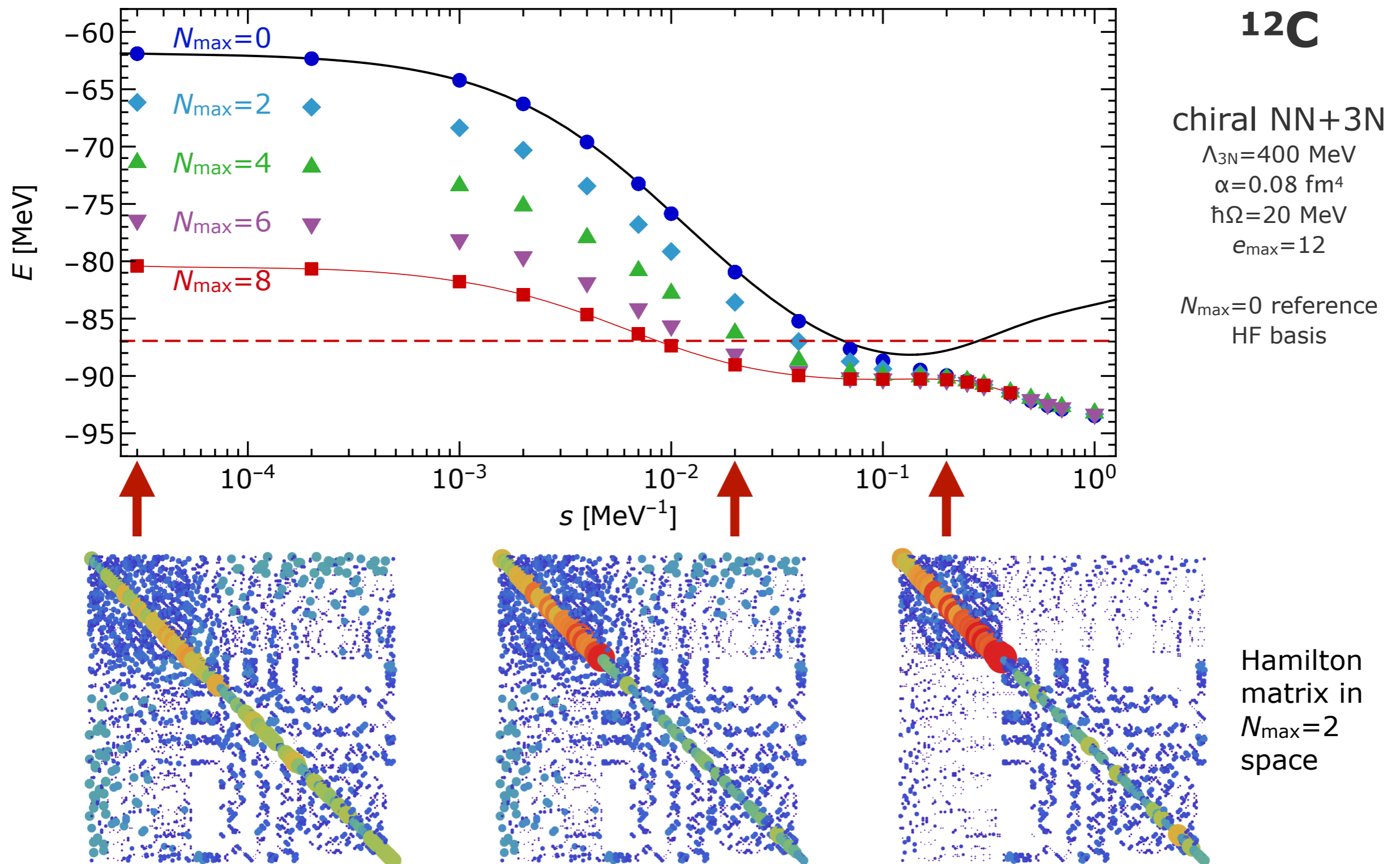
- IM-SRG evolution of multi-reference normal-ordered Hamiltonian and other operators
- decoupling of particle-hole excitations, i.e., pre-diagonalization in many-body space

**NCSM**  
many-body solution

- use in-medium evolved Hamiltonian for a subsequent NCSM calculation
- access to ground and excited states and full suite of observables

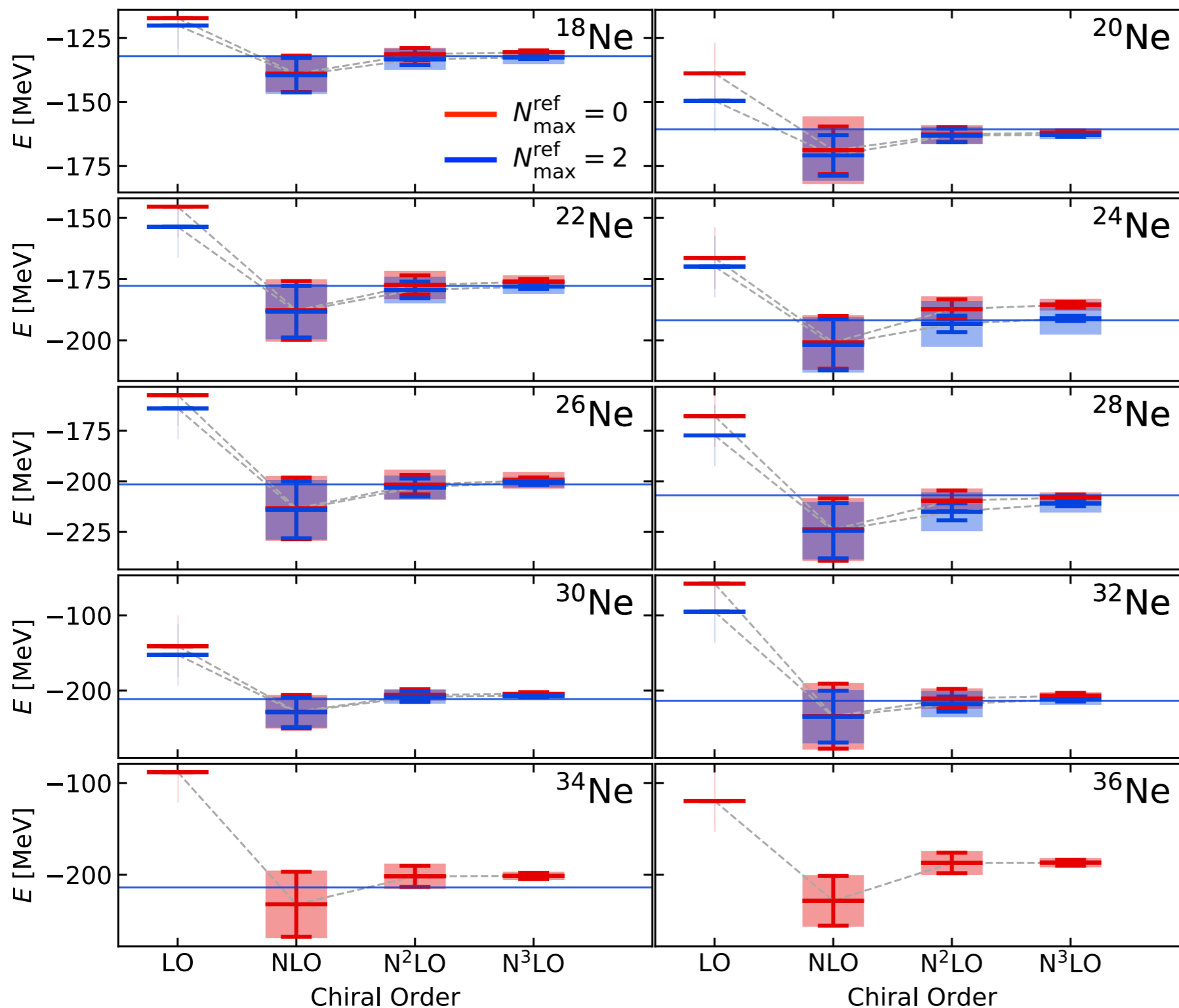
# In-Medium NCSM: Flow Evolution

Gebrerufael, Vobig, Hergert, Roth; PRL 118, 152503 (2017)



# Neon Isotopes: Ground-State Energies

Mongelli et al., in preparation



- amazing reproduction of experimental energies for all isotopes
- uncertainties under control

family of non-local NN+3N interactions  
*T. Hüther et al., PLB 808, 135651 (2020)*

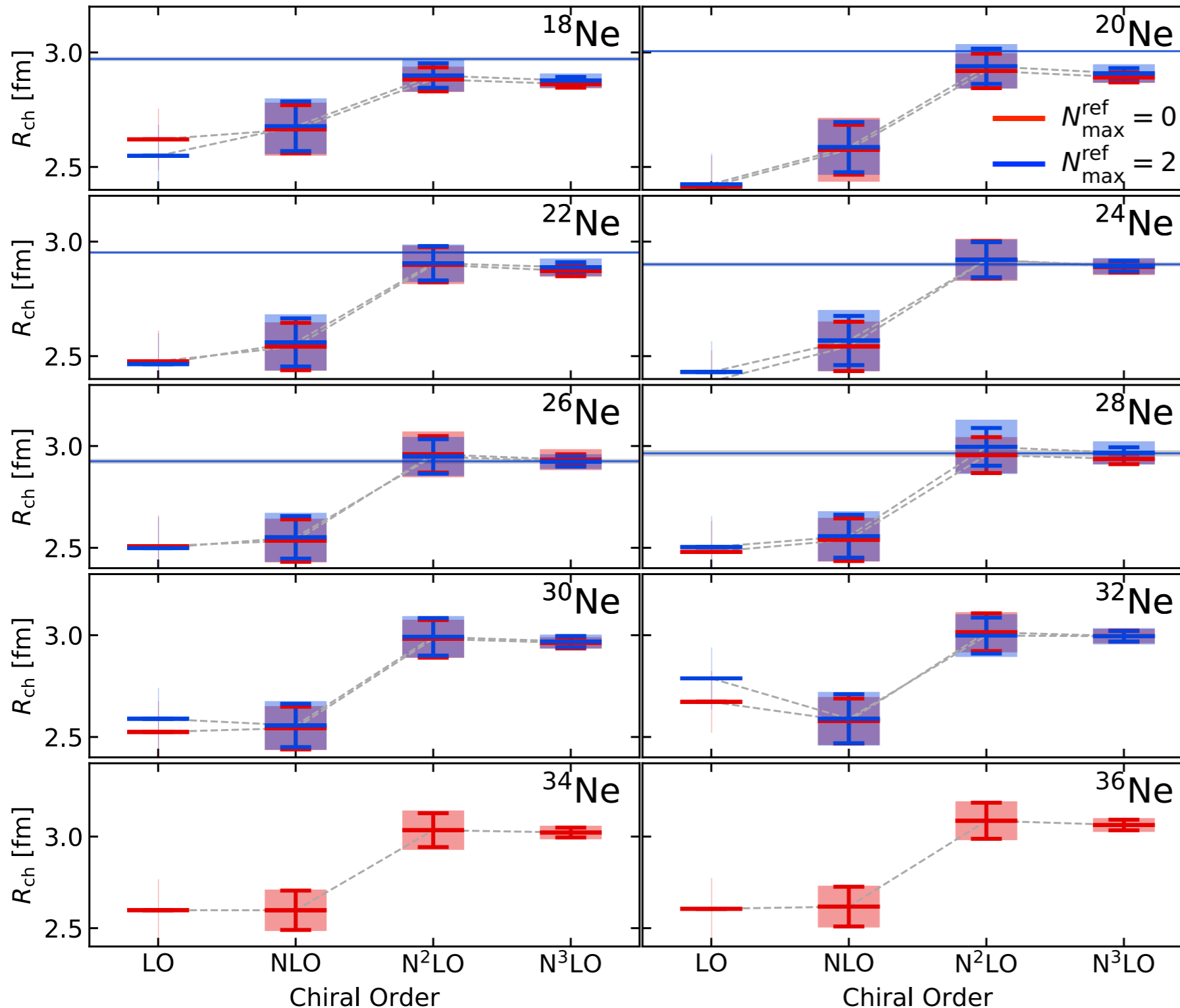
$\Lambda = 500$  MeV  
 $\alpha = 0.04$  fm<sup>4</sup>  
 $\hbar\Omega = 20$  MeV  
 $e_{\max} = 12$   
 NAT basis  
 $N_{\max}^{\text{ref}} = 0, 2$   
 $N_{\max} = 4$

*error bars:*  
 68% interaction  
 uncertainties

*error bands:*  
 interaction +  
 many-body  
 uncertainties

# Neon Isotopes: Charge Radii

Mongelli et al., in preparation



- excellent description of radii, slight underestimation for light isotopes
- stable results in N<sup>2</sup>LO and N<sup>3</sup>LO

family of non-local NN+3N interactions  
*T. Hüther et al., PLB 808, 135651 (2020)*

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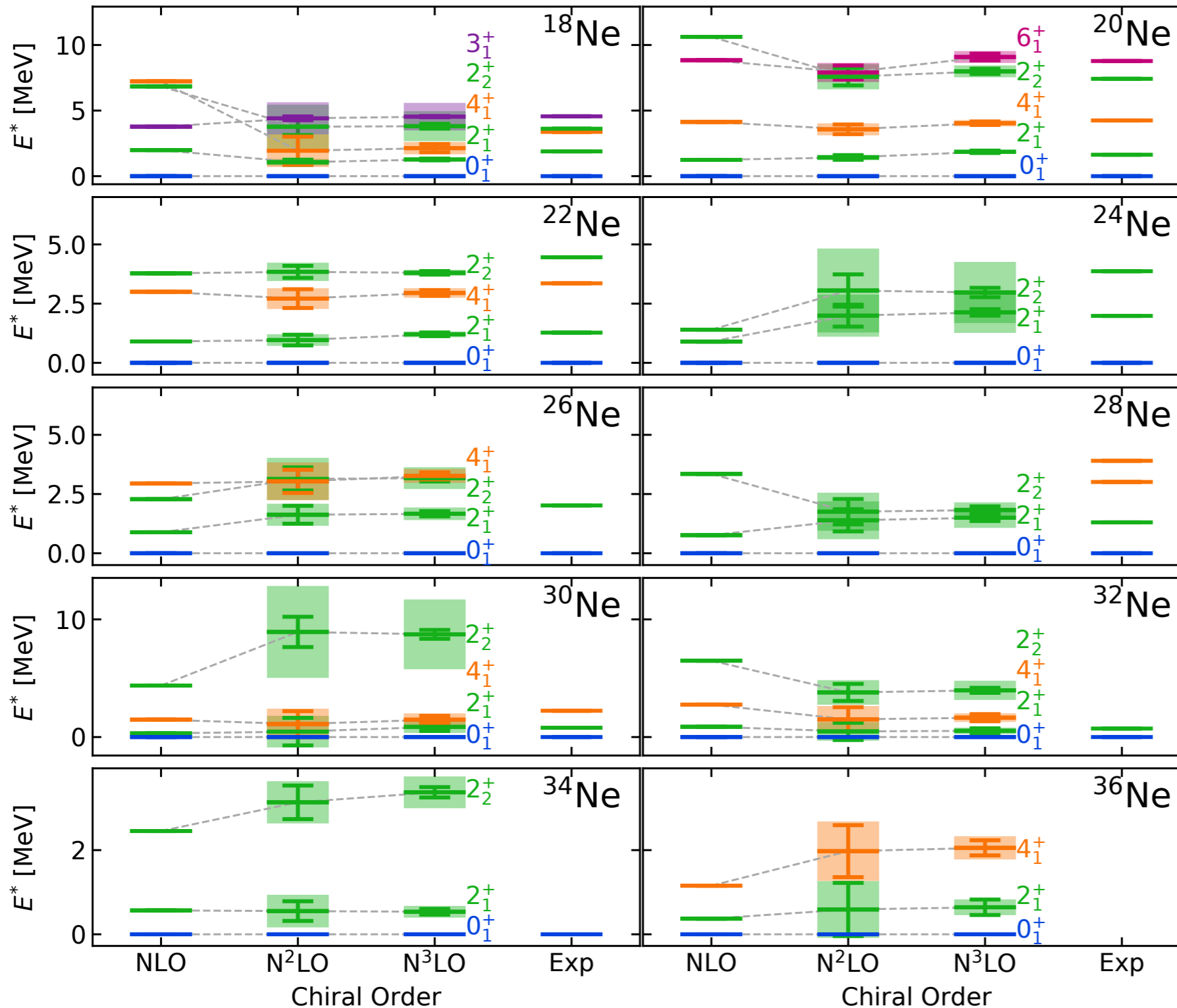
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*error bands:*  
 interaction +  
 many-body  
 uncertainties



# Neon Isotopes: Excitation Energies

Mongelli et al., in preparation



■ excellent description of excitation spectra

family of non-local NN+3N interactions  
*T. Hüther et al., PLB 808, 135651 (2020)*

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error bars:  
 68% interaction  
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$e_{\max} = 12$   
 NAT basis

$N_{\max}^{\text{ref}} = 2$   
 $N_{\max} = 4$

error bands:  
 interaction +  
 many-body  
 uncertainties

# Following Talks

## A02

Sebastian Dietz

-

*Energy distribution of  
 $^3n$  systems*

Tobias Wolfgruber

-

*From chiral interactions to  
NCSM observables*

Julius Müller

-

*First results from the  
Gamow-NCSM*

## A04

Takayuki Miyagi

-

*Heavy-mass frontier in  
nuclear ab initio calculations*

## B05

Jonas Keller

-

*Nuclear EOS for arbitrary  
proton fraction and  
temperature*

Andreas Geissel

-

*Towards the EOS of  
neutron stars*