

Exploration of extremes of nuclei at SAMURAI at RIBF

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- SAMURAI Facility at RIBF
- Recent Activities at SAMURAI with NeuLAND
- Coulomb Breakup of Halo Nucleus ^{31}Ne
- Spectroscopy of heavy oxygen isotopes ($^{26,27,28}\text{O}$)
- Summary and Perspectives

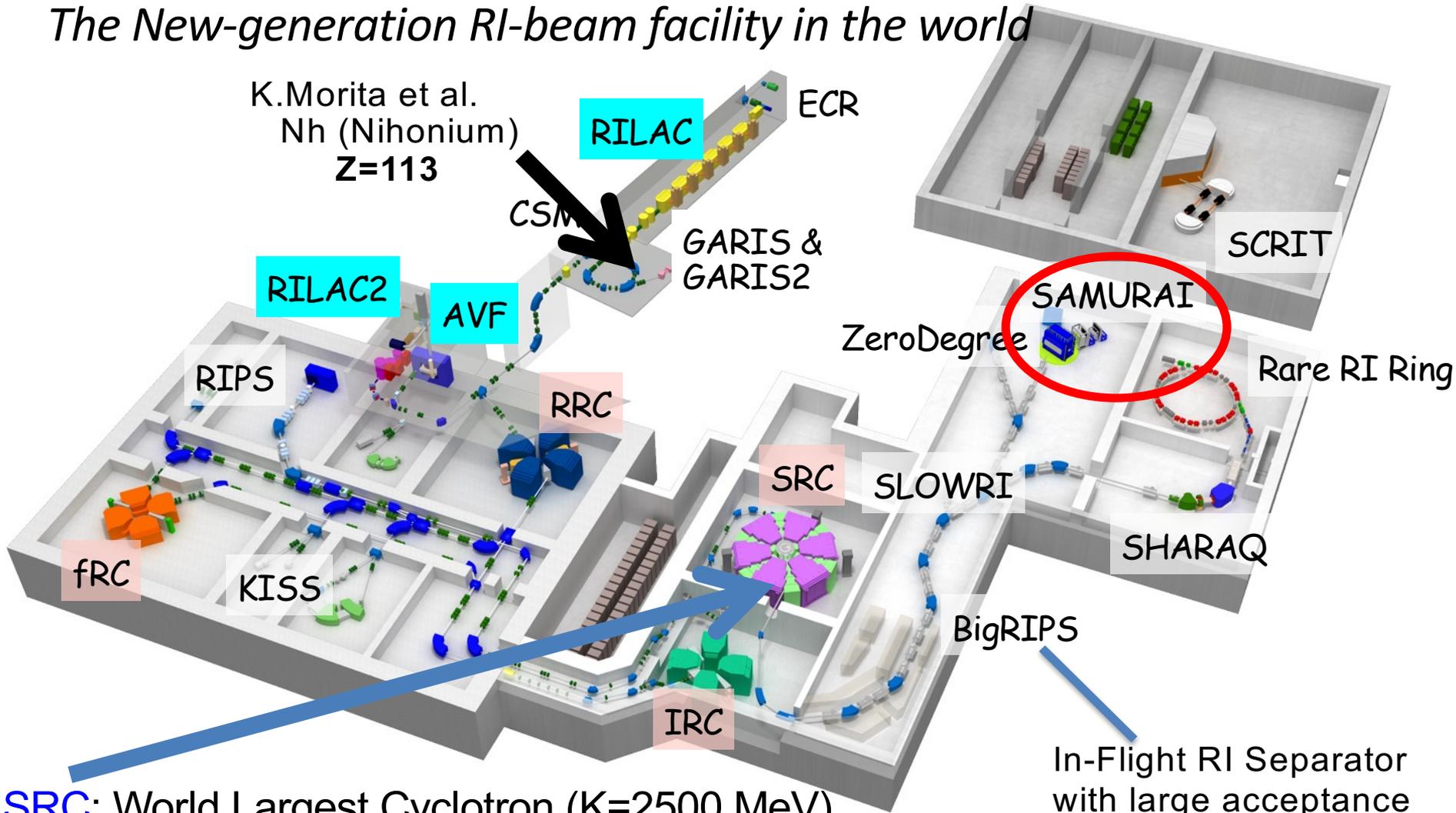


RI Beam Factory at RIKEN

2007~

The New-generation RI-beam facility in the world

K.Morita et al.
Nh (Nihonium)
Z=113



SRC: World Largest Cyclotron (K=2500 MeV)

Heavy Ion Beams up to ^{238}U at 345MeV/u

eg. ^{48}Ca : $\sim 700\text{pnA}$ ($\sim 4 \times 10^{12}$ pps) ~ 10 times compared to 2008

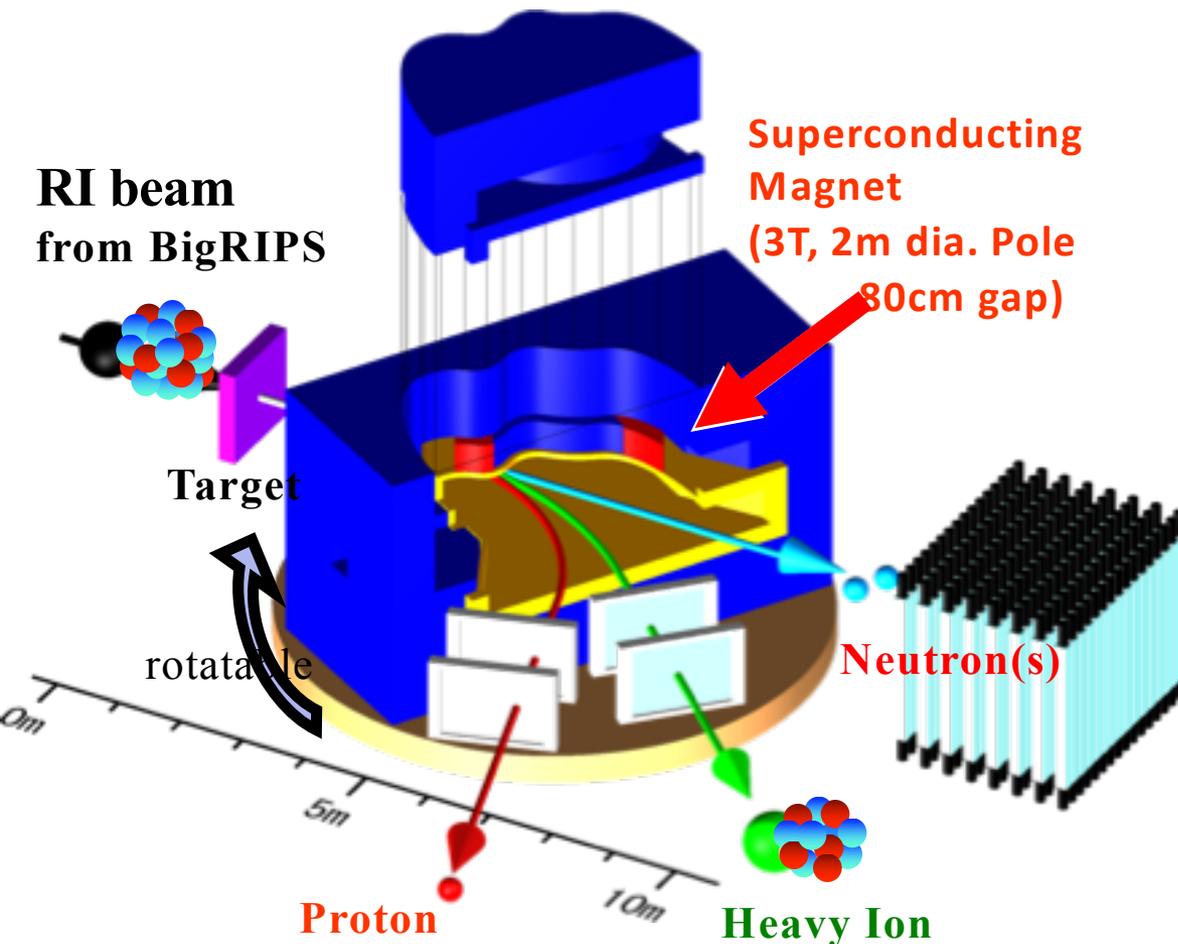
^{238}U : $\sim 30\text{pnA}$ ($\sim 2 \times 10^{11}$ pps) $\sim 10^3$ times compared to 2007

In-Flight RI Separator
with large acceptance

SAMURAI at RIBF/RIKEN

Superconducting **A**nalyzer for **M**ulti-particle from **R**adio **I**sotope Beam

Kinematically Complete measurements by detecting multiple particles in coincidence



Large momentum acceptance

$$B\rho_{\max} / B\rho_{\min} \sim 2 - 3$$

Good Momentum Resolution

$$\Delta p/p \sim 1/1000$$

$$\rightarrow A/\Delta A > 100 (>5\sigma)$$

Large Bending Angle ($\sim 60\text{deg}$)

+4 Tracking Detectors

T.Kobayashi NIMB **317**, 294 (2013)

Large angular acceptance for n

$$+8.8 \text{ deg (H)} \times +4.4 \text{ deg (V)}$$

$$(\sim 50\% \text{ coverage } < E_{\text{rel}} \sim 5 \text{ MeV})$$

TN, Y.Kondo, NIMB **376**, 156 (2016).

Moderate Erel Resolution

$$\Delta E = 200 \text{ keV } (\sigma) \text{ at } E_{\text{rel}} = 1 \text{ MeV}$$

Stage: Rotatable (-5 -- 95 degrees)

\rightarrow Variety of Physics Opportunities

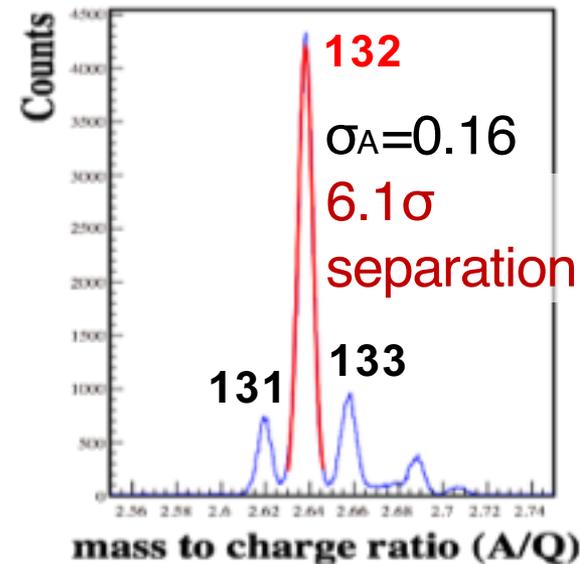
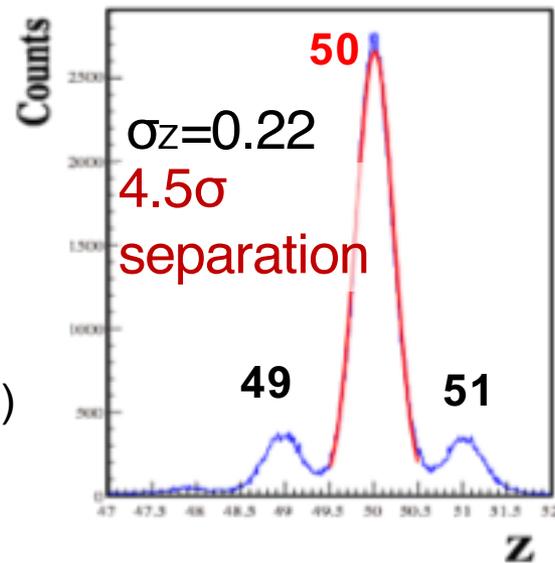
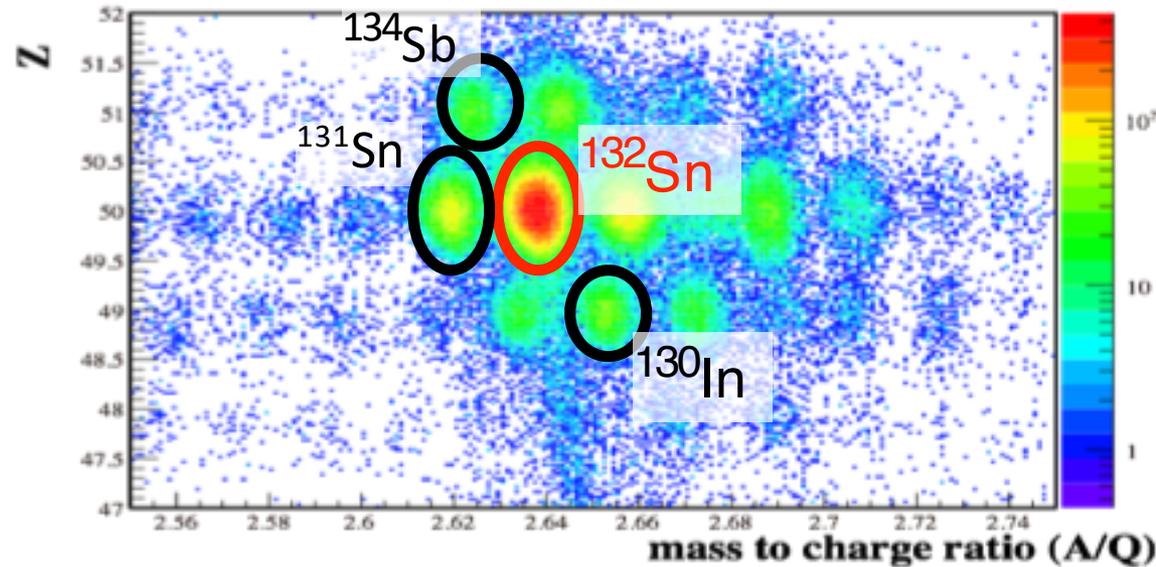
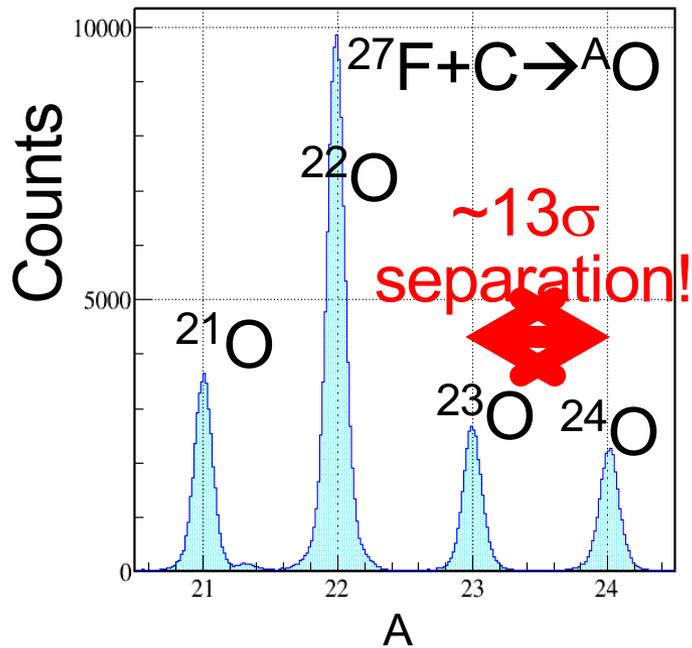
SAMURAI Since 2012

Superconducting Analyzer for Multi-particle from Radio Isotope Beam



March 2012

PID at SAMURAI : Mass Resolution

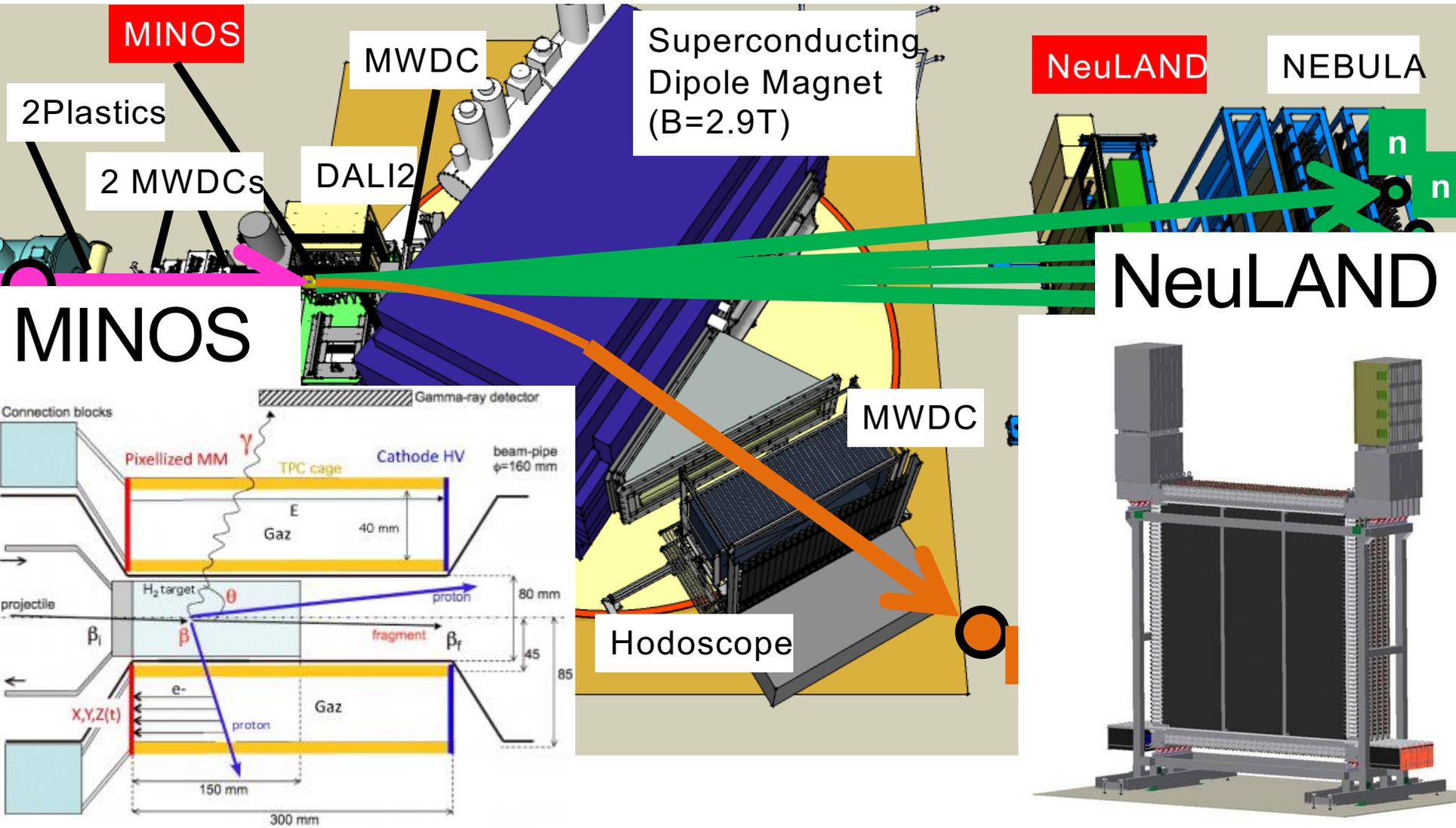


Clear Particle identification
 → High resolving power

^{24}O : $A/\Delta A = 130$ (FWHM), $62(5\sigma)$
 ^{132}Sn : $A/\Delta A = 340$ (FWHM), $160(5\sigma)$
 $Z/\Delta Z = 96$ (FWHM), $45(5\sigma)$

Great Contributions
 from **R3B (2015-17, NEULAND)**
 and **CEA Saclay (2014-2017, MINOS)**

Slide by Y.Kondo



Experiments with NEULAND (2015-17)

(Partial) List of the experiments using NEULAND+SAMURAI:

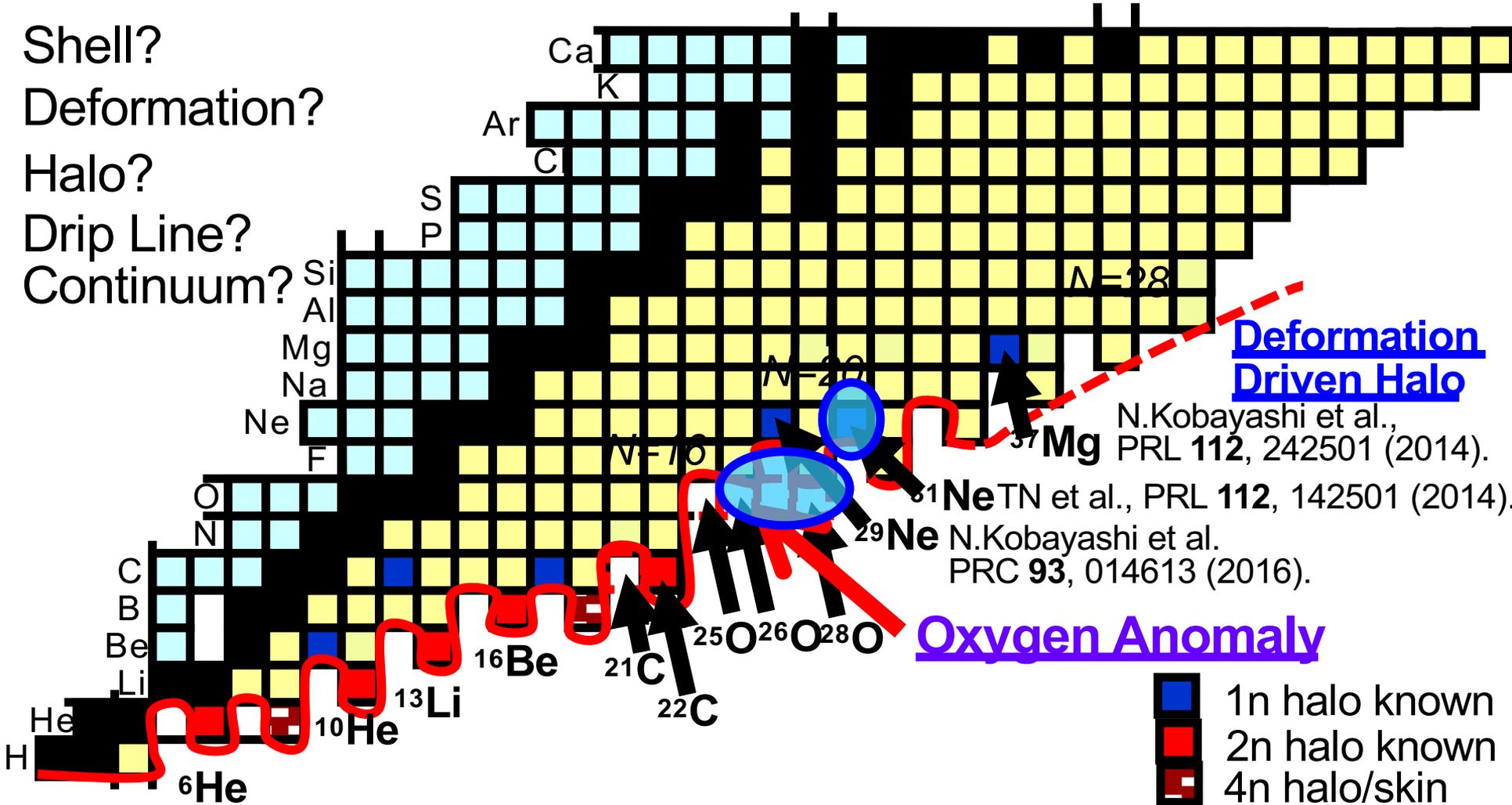
- ✓ Spectroscopy of unbound oxygen isotopes II (Y.Kondo) [27,28O](#)
- ✓ Study of density dependence of the symmetry energy with the measurements of charged pion ratio in heavy RI collisions (W.G. Lynch, T. Murakami, T. Isobe, B. Tsang)
- ✓ Search for ^{22}C (2^+), ^{21}B , ^{23}C and ^{25}N : Structure at and beyond the $N=16$ sub-shell closure (N.A. Orr)
- ✓ Spectroscopy of Odd-A Nuclei in the Island of Inversion (N.Kobayashi) [31Ne etc.](#)
- ✓ Measurement of the neutron-decay lifetime of the ^{26}O ground state at the SAMURAI setup at RIBF (C. Caesar)
- ✓ Electric dipole response of neutron-rich Ca isotopes (T. Kobayashi, Y.Togano)
- ✓ SEASTAR: 3rd Campaign (including "Spectroscopy of neutron-rich Ca isotopes" by J. Lee) (A.Obertelli, P.Dornenbal)
- ✓ Many-neutron systems: search for superheavy ^7H and its tetra-neutron decay (K.Kisamori, M.Miguel)
- ✓ Dipole response of the drip-line nuclei ^8He , ^{24}O , and ^{29}F Search for the genuine soft dipole mode (T.Aumann, T.Nakamura)
- ✓ Investigation of the $4n$ system at SAMURAI by measuring $(p,p\alpha)$ quasi-free scattering at large momentum transfer in complete kinematics (S. Paschalis, D. Rossi)

Evolution Towards the Stability Limit

Where is the neutron drip line?

What are characteristic features of drip-line nuclei?

How does nuclear structure evolve towards the drip line?

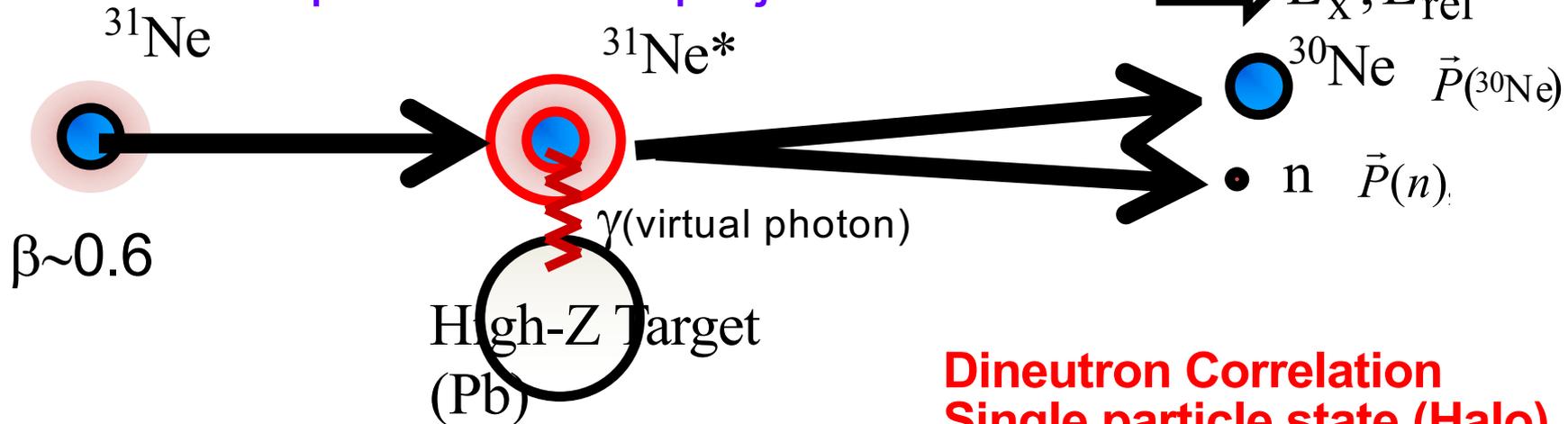




Reaction Probes

Coulomb Breakup

→ Photon absorption of a fast projectile



Equivalent Photon Method

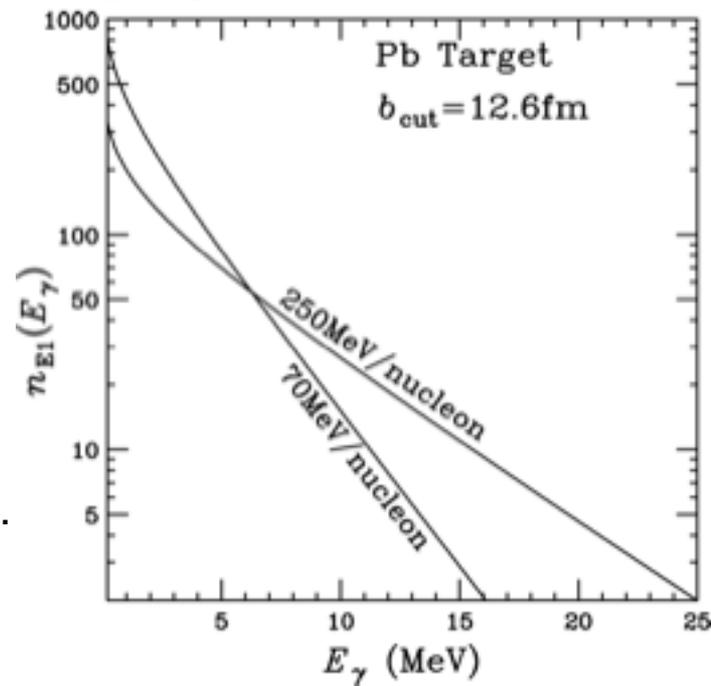
$$\frac{d\sigma_{CB}}{dE_x} = \frac{16\pi^3}{9\hbar c} N_{E1}(E_x) \frac{dB(E1)}{dE_x}$$

Cross section = (Photon Number) x (Transition Probability)

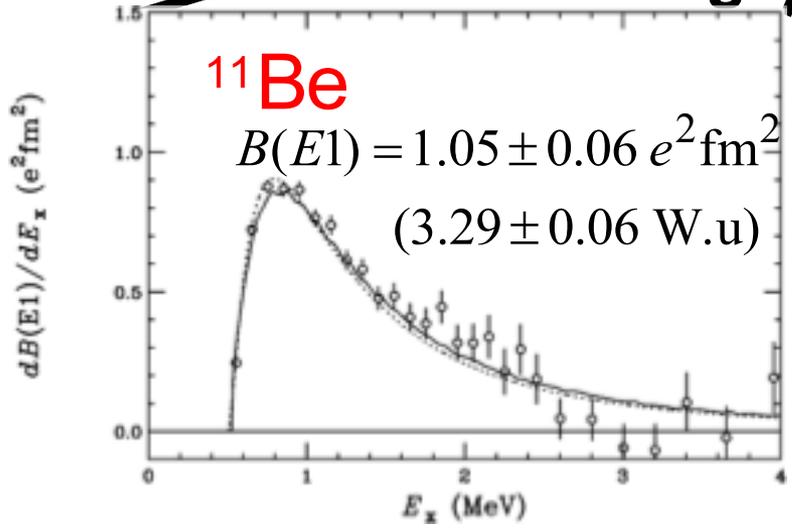
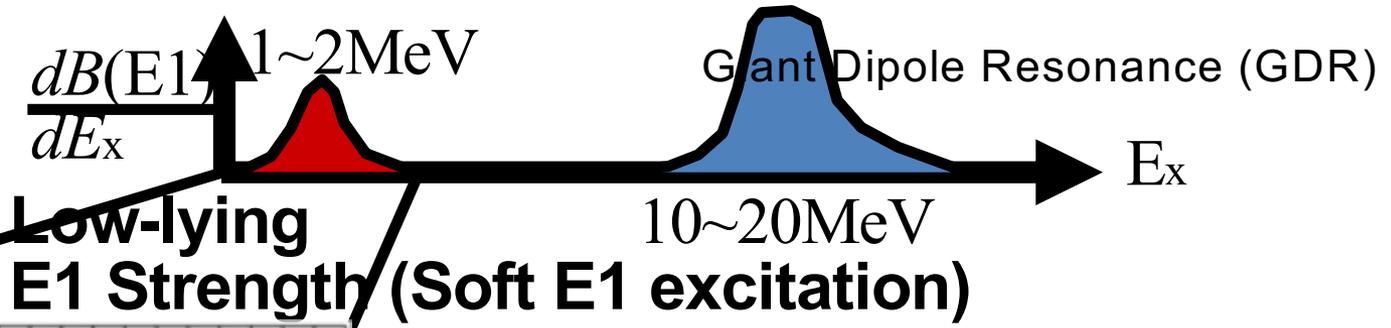
C.A. Bertulani, G. Baur, Phys. Rep. **163**, 299(1988).
T. Aumann, T. Nakamura, Phys. Scr. T**152**, 014142(2013).

Halo → Soft E1 Excitation
(E1 Concentration at $E_x < 1\text{MeV}$)

Dineutron Correlation
Single particle state (Halo)



Coulomb Breakup and E1 Response--Case of 1n Halo



N.Fukuda *et al.*, PRC70, 054606 (2004).
 T.Nakamura *et al.*, PLB 331,296(1994).
 Palit *et al.*, PRC68, 034318(2003).
 T.Aumann, T.Nakamura, Phys. Scr. T152 (2013) 014012.



Direct Breakup Mechanism

$$\rho \sim C^2S |\exp(-r/\lambda)/r|^2$$

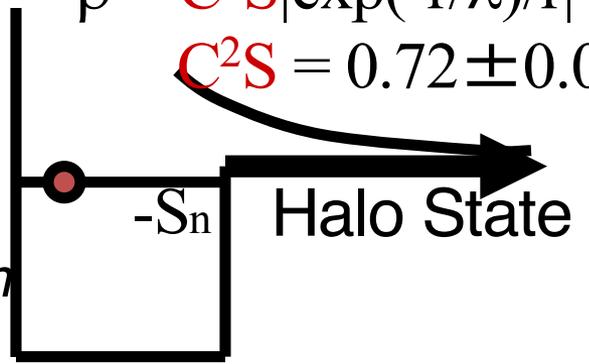
$$C^2S = 0.72 \pm 0.04$$

E1 Strength

$$\frac{dB(E1)}{dE_x} \propto \left| \langle \exp(iqr) | \frac{Z}{A} r Y^1_m | \Phi_{gs} \rangle \right|^2$$

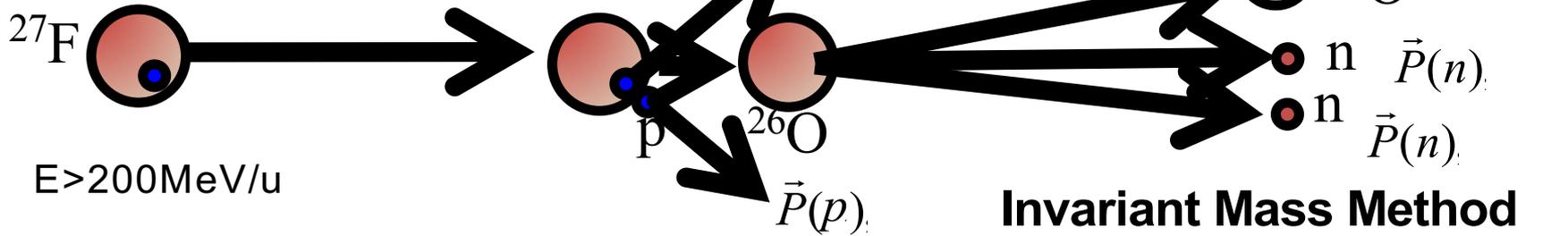
$$\propto C^2S \left| \langle \exp(iqr) | \frac{Z}{A} r Y^1_m | S_{1/2} \rangle \right|^2$$

Fourier Transform

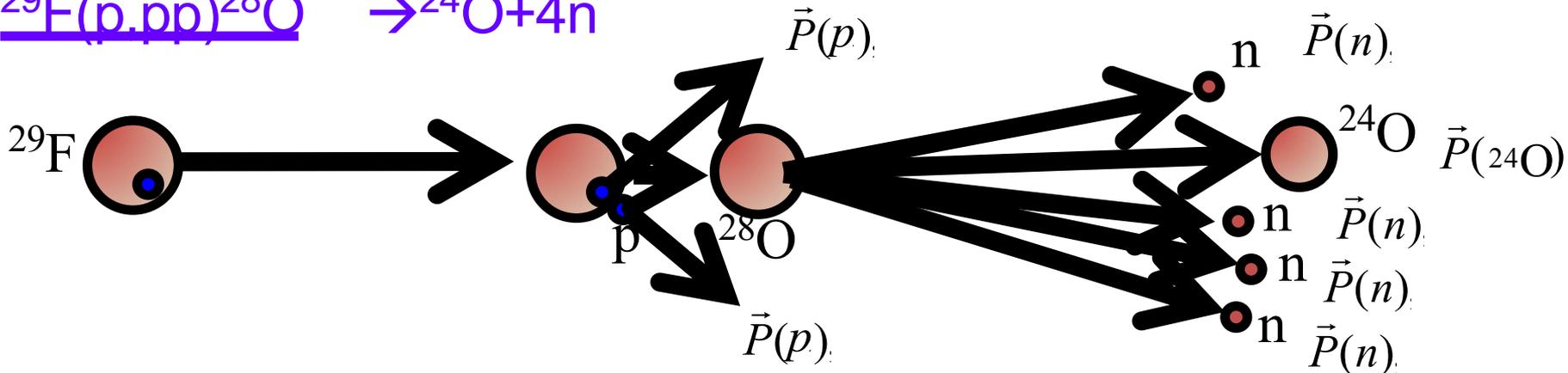


Soft E1 Excitation of 1n halo—Sensitive to S_n, l, C^2S

Quasi-free Scattering -with neutron decay



Missing Mass Method



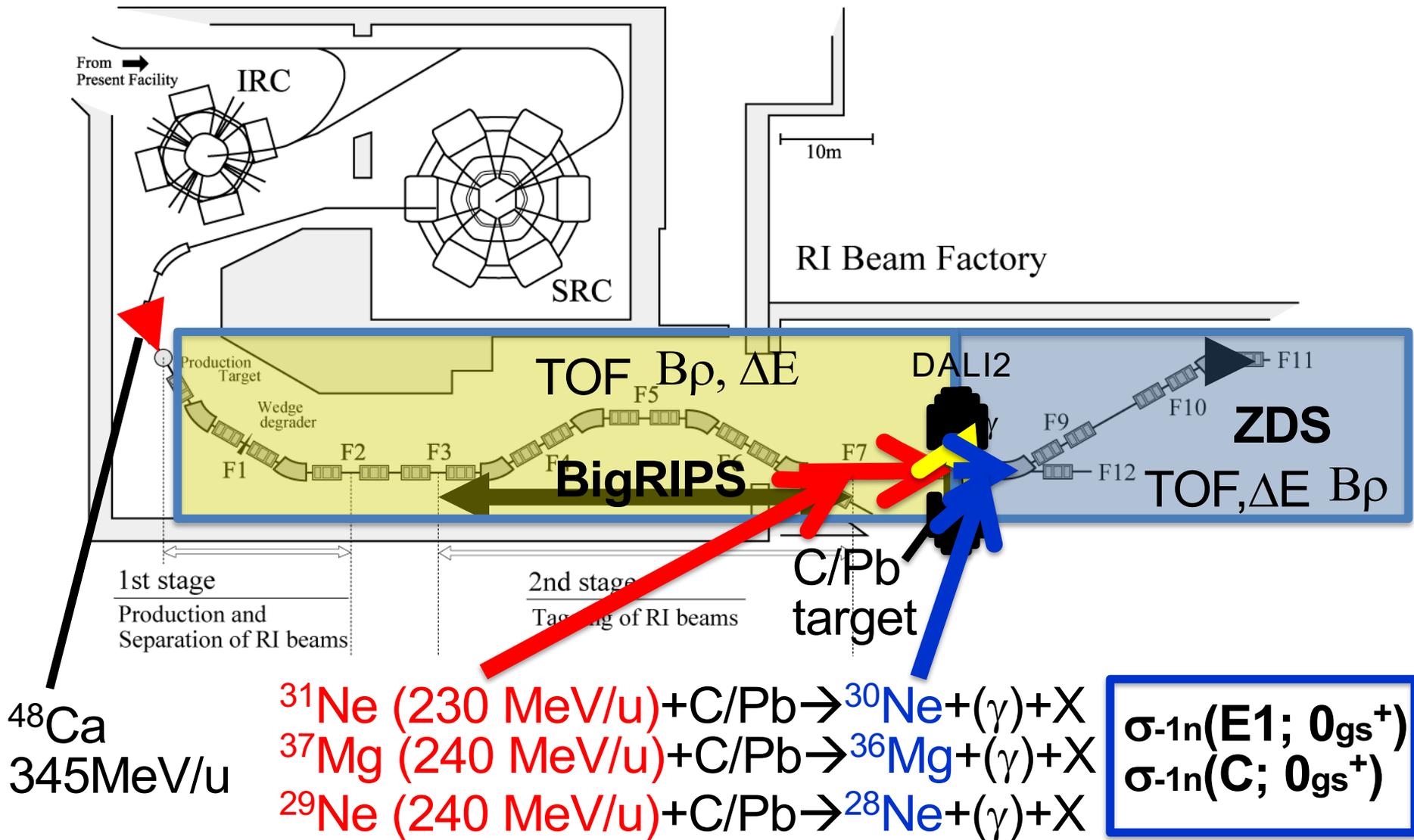
Invariant Mass Method: + High Yield, + Good Resolution ~ a few 100 keV
 - Require Measurement of All the Decay Particles

Missing Mass Method: - Low Yield, - Worse Resolution ~ a few MeV
 + Measurement of projectile and recoil protons only

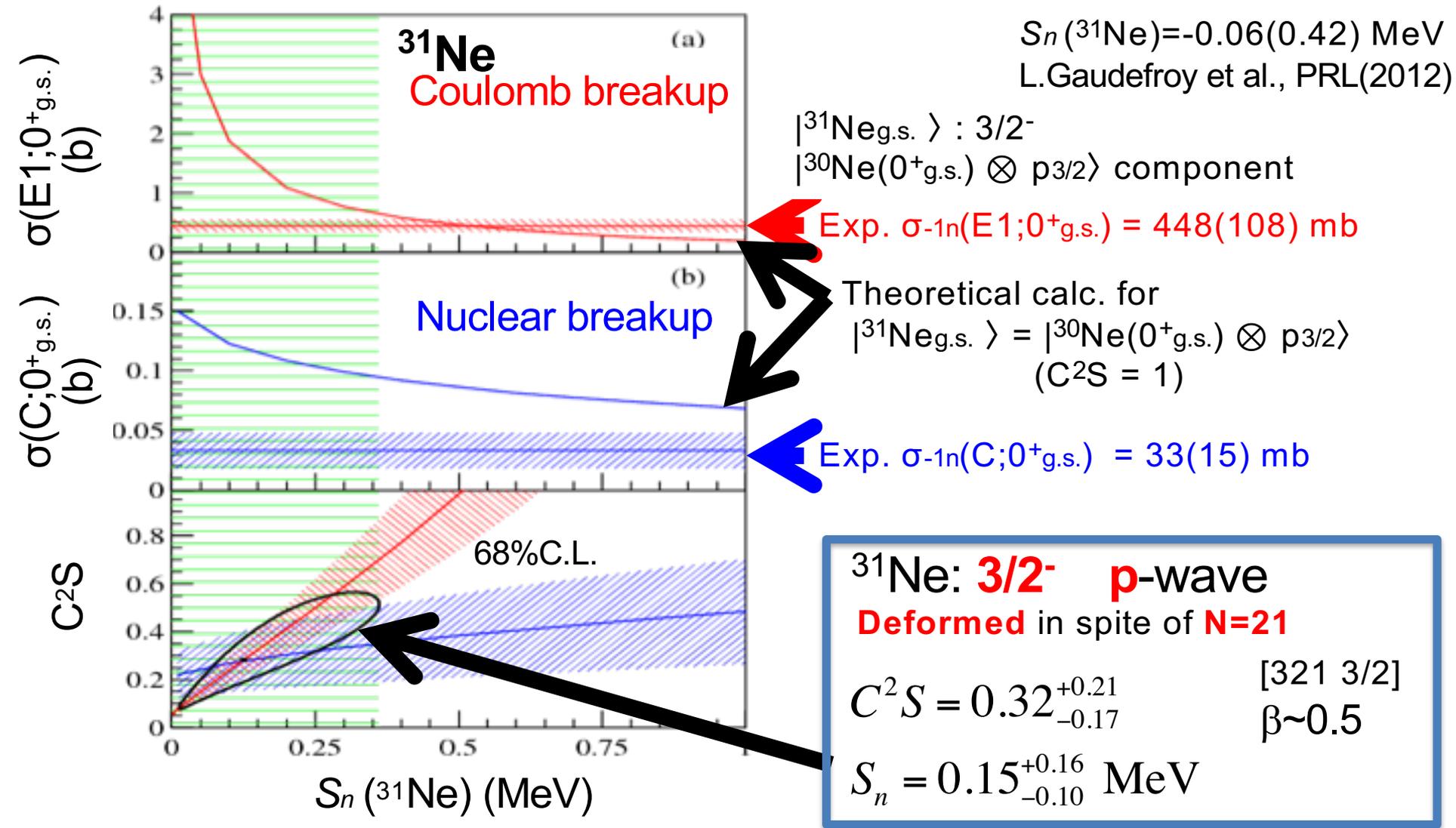
 Breakup of 1n Halo Nuclei
in Island of Inversion (^{29}Ne , ^{31}Ne , ^{37}Mg)

N. Kobayashi,
T. Tomai,
Y. Kondo,
TN et al.

Inclusive nuclear/Coulomb Breakup at BigRIPS & ZDS at RIBF



Deformation Driven p-wave Halo --- ^{31}Ne , ^{37}Mg , ^{29}Ne

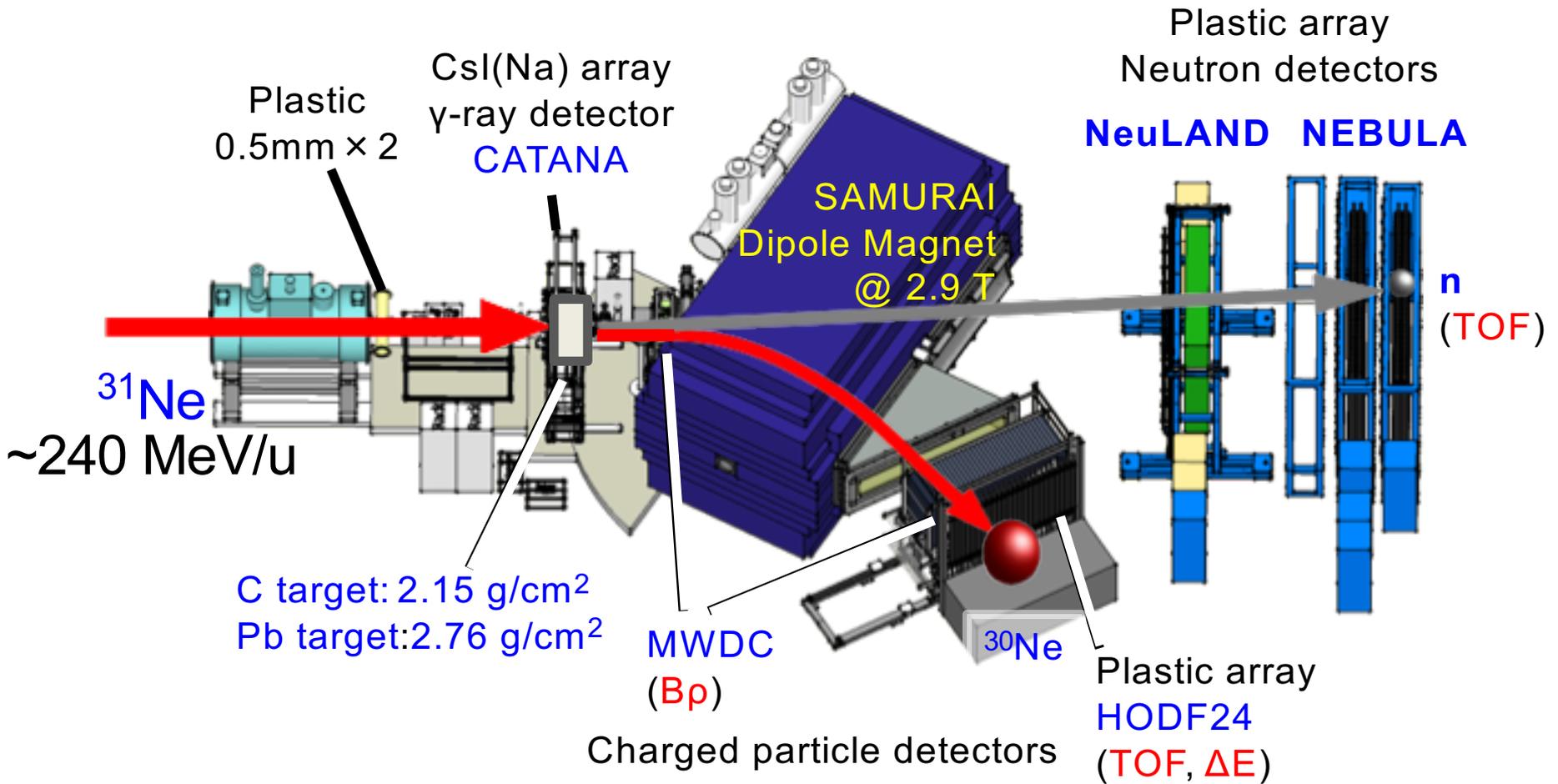


^{31}Ne : TN, N.Kobayashi et al., PRL **112**, 142501 (2014). $3/2^-$ $S_n=150(16)$ keV
 ^{37}Mg : N.Kobayashi, TN et al., PRL **112**, 242501 (2014). $3/2^-/1/2^-$ $S_n=220(12)$ keV
 ^{29}Ne : N.Kobayashi, TN et al., PRC **93**, 014613 (2016). $3/2^-$ $S_n=960(140)$ keV

Inclusive at BigRIPS/ZDS
→ Exclusive at SAMURAI

SAMURAI Experiment

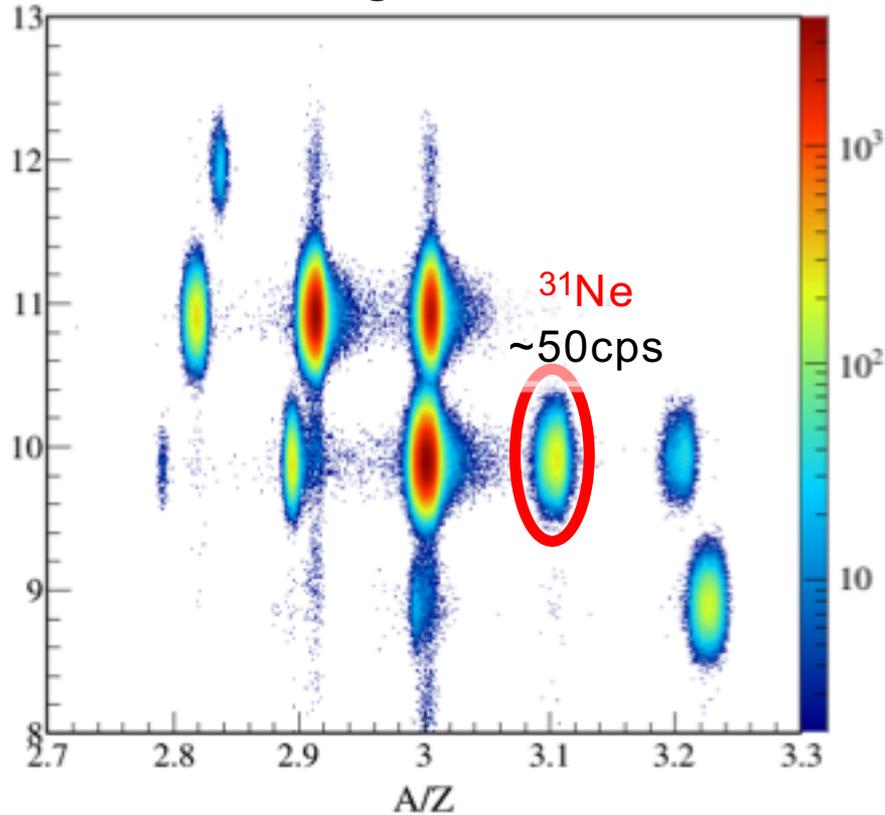
Full Exclusive Coulomb Breakup Measurement of ^{31}Ne T.Tomai et al.
-Autumn, 2015



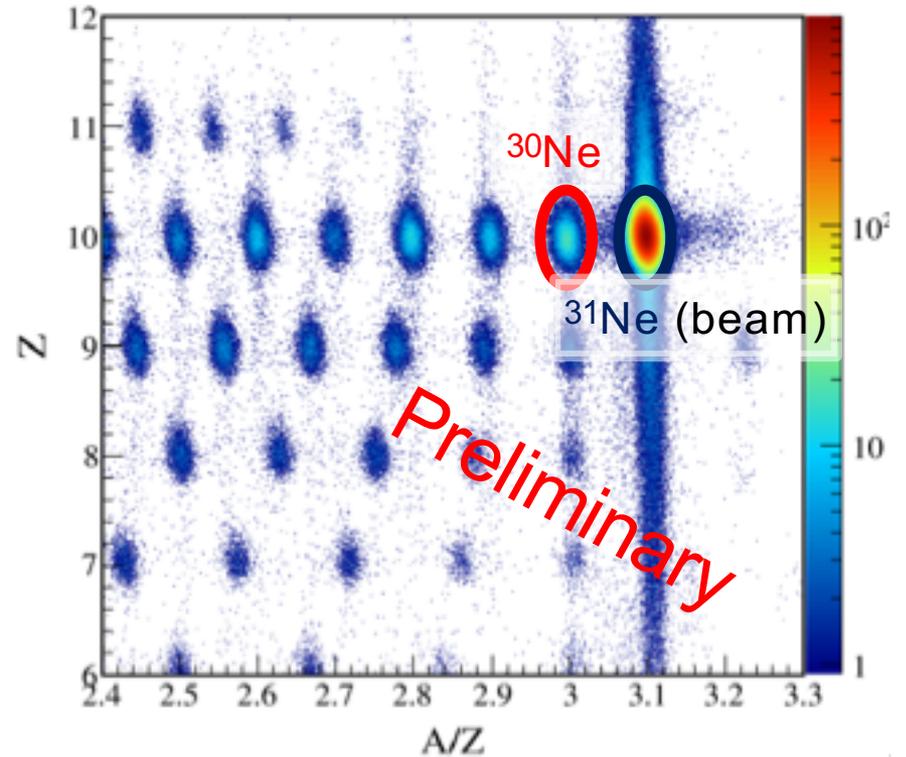
PID Spectra

T.Tomai et al.

Incoming RI Beam



Outgoing Fragments from ^{31}Ne (C tgt)

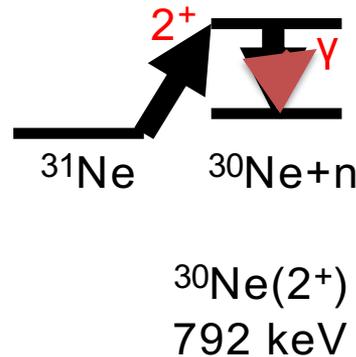


Intensity: **10 times higher**
than previous RIBF experiment in 2008

γ -ray spectrum : Excited state component of Coulomb breakup reaction

Preliminary

- How is the excited state component of $^{30}\text{Ne}(2^+)$?

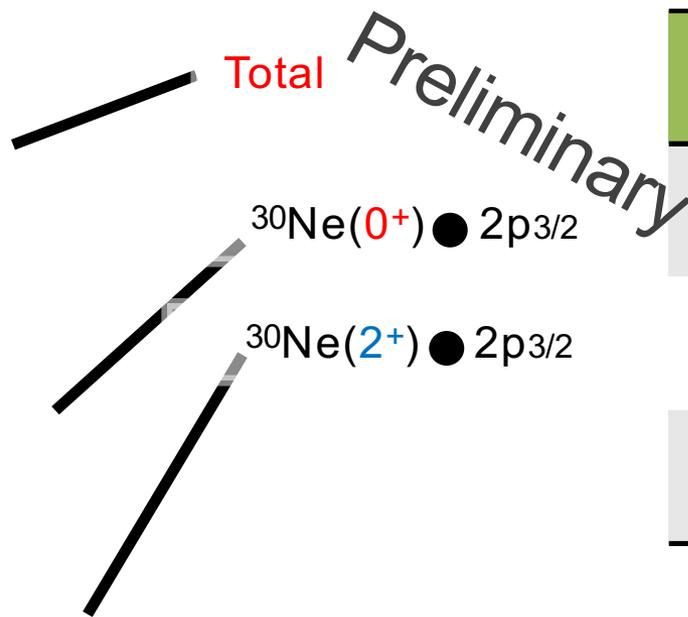


$$|^{31}\text{Ne}(3/2^-)\rangle = \alpha |^{30}\text{Ne}(0^+) \otimes 2p_{3/2}\rangle + \beta |^{30}\text{Ne}(2^+) \otimes 2p_{3/2}\rangle + \dots$$

Pb target	Exclusive (mb) (require n) Integral $E_{\text{rel}}=0-5\text{MeV}$
$^{31}\text{Ne} \rightarrow ^{30}\text{Ne}(\text{total})$	479(18)
$^{31}\text{Ne} \rightarrow ^{30}\text{Ne}(2^+)$	136(24)
$^{31}\text{Ne} \rightarrow ^{30}\text{Ne}(0^+)$	343(30)
Ratio($0^+ : 2^+$)	72(6)% : 28(5)%

Coulomb breakup of ^{31}Ne : Energy Spectrum

Tomai et al.



	S_n (MeV)	C^2S $^{30}\text{Ne}(0^+);3/2^-$	C^2S $^{30}\text{Ne}(2^+);3/2^-$
This work	0.30(1) (preliminary, statistical error only)	0.32(1)	0.41(2)
Prev. work*	$0.15^{+0.16}_{-0.10}$	$0.32^{+0.21}_{-0.17}$	---
SDPF-M		0.21	0.34

*TN, N. Kobayashi et al., PRL112, 142501 (2014)

Consistent with previous work

- Fitted with 2 components

$$\alpha |^{30}\text{Ne}(0^+) \bullet 2p_{3/2} \rangle + \beta |^{30}\text{Ne}(2^+) \bullet 2p_{3/2} \rangle$$

72%

28%

γ -ray data

Main Halo Component

Sub-Halo Component

→ **Doubly-component Halo?**

$$|^{30}\text{Ne}(0^+) \bullet 2p_{3/2} \rangle : |^{30}\text{Ne}(2^+) \bullet 2p_{3/2} \rangle$$

→ Correspondence with Nilsson/Particle-Rotor Model → Deformation Parameter?

c.f. Y.Urata, K.Hagino, H.Sagawa, PRC83, 041303(R) (2011).

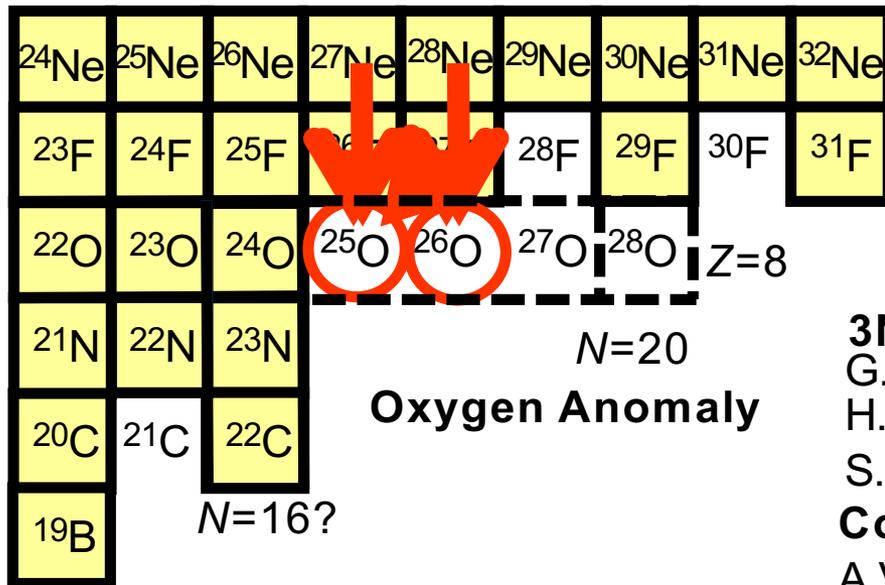
 Spectroscopy
of Barely Unbound $^{26,27,28}\text{O}$
(& Other studies on unbound oxygen isotopes)

[Yosuke Kondo
et al.](#)

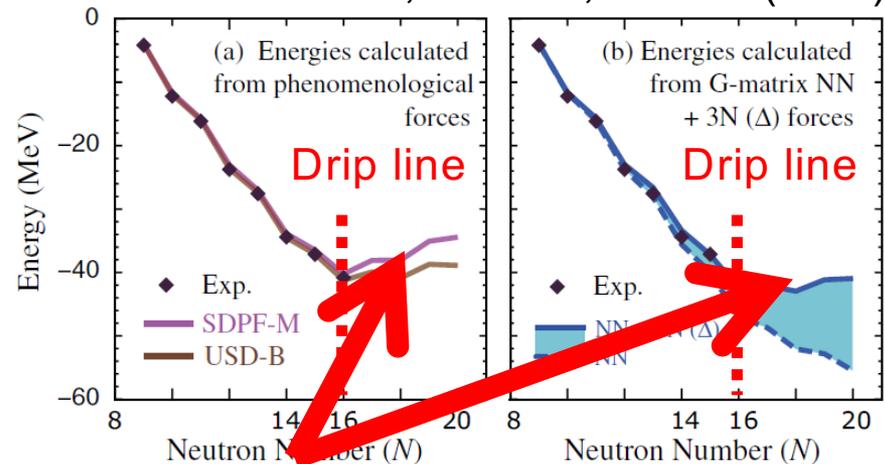
Study of unbound nuclei ^{25}O and ^{26}O at SAMURAI

Spokesperson Yosuke Kondo

Experimental study of unbound oxygen isotopes towards the possible double magic nucleus ^{28}O



T. Otsuka et al., PRL105, 032501 (2010).



3N force: significant at $N > 16$

G. Hagen et al., PRL108, 242501(2012).

H. Hergert et al., PRL110, 242501(2013).

S.K.Bogner et al., PRL113, 142501(2014).

Continuum Effect:

A.Volya, V.Zelevinski, PRL94,052501(2005).

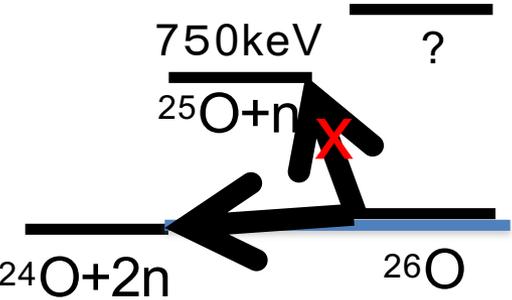
K. Tsukiyama, T. Otsuka, PTEP2015, 093D01 (2015).

nn correlations:

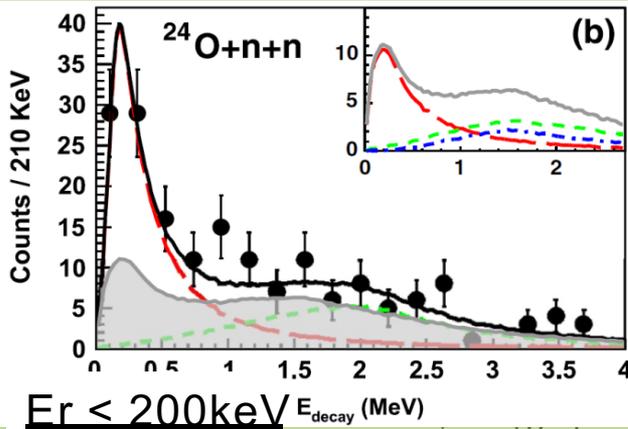
L.V. Grigorenko et al., PRL111,042501(2013).

K. Hagino, H. Sagawa PRC89,014331(2014).

2n radioactivity of ^{26}O ?

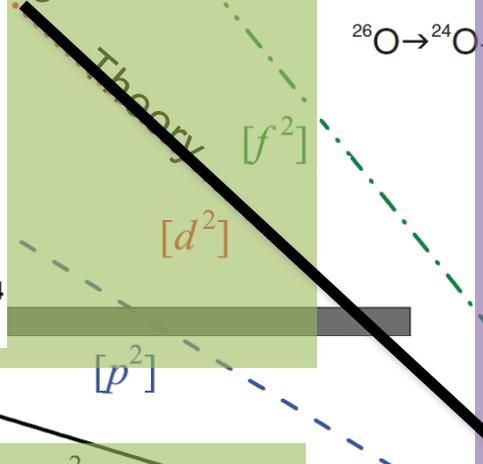


E. Lunderberg et al.
 PRL108, 142503 (2012)

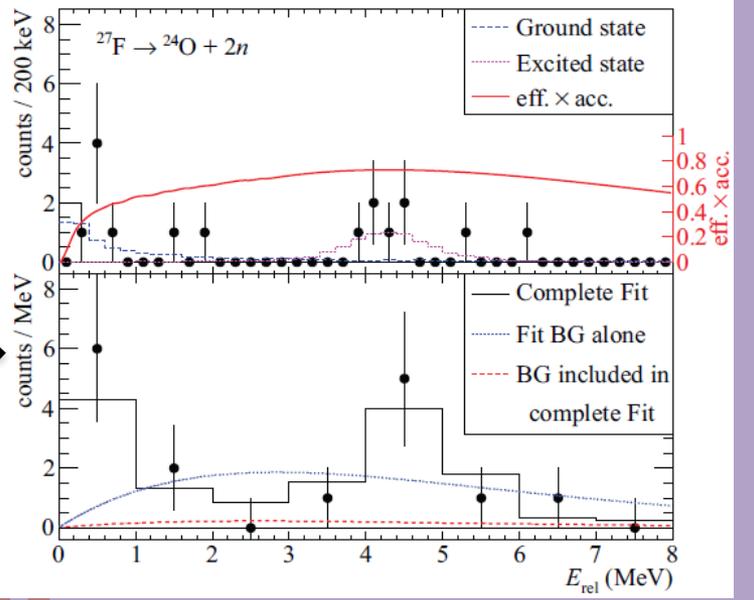


$^{26}\text{O}: ^{24}\text{O}(0^+) \otimes (vd_{3/2})^2$

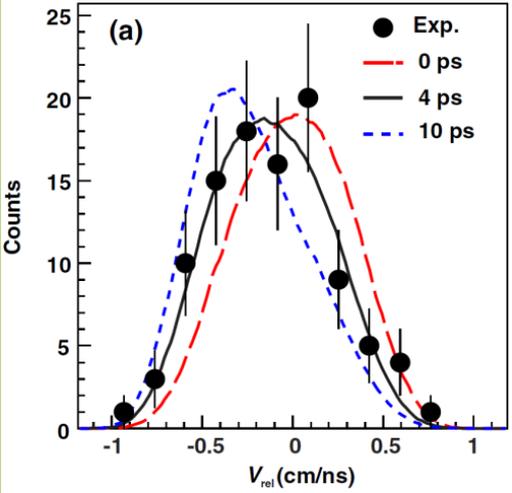
rigorenko et al. PRC 84, 021303 (2011)



C. Caesar et al. PRC88, 034313 (2013)



Z. Kohley et al, PRL110, 152501 (2013)



$T_{1/2} = 4.5^{+1.1}_{-1.5} \text{ ns}$
 (3ps systematic error)

\rightarrow 2n radioactivity?

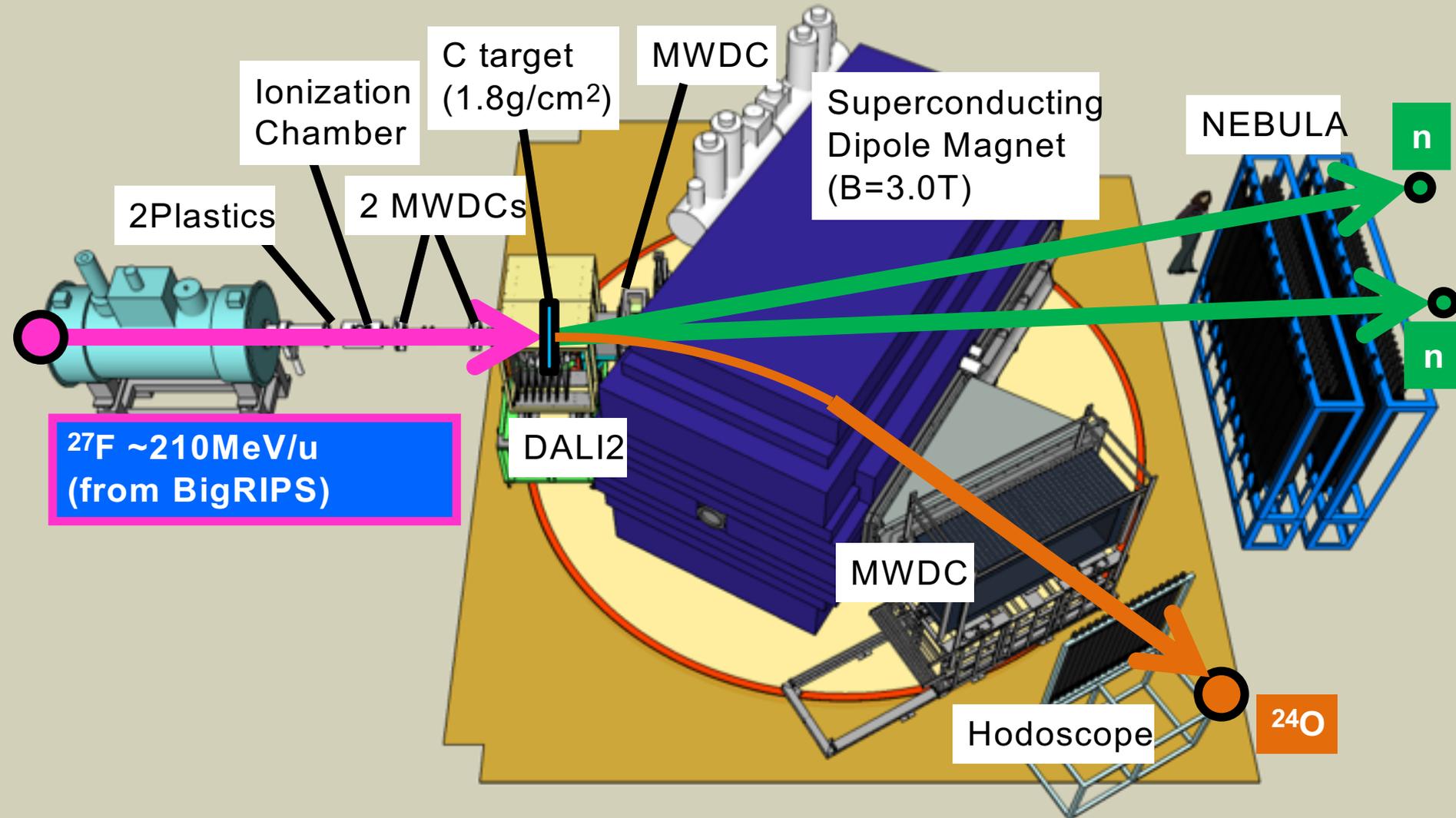
Usual 1n decay
 $\Gamma \sim \text{MeV or keV}$

$E_r < 120 \text{ keV}$ (95% CL)
 $\tau < 5.7 \text{ ns}$

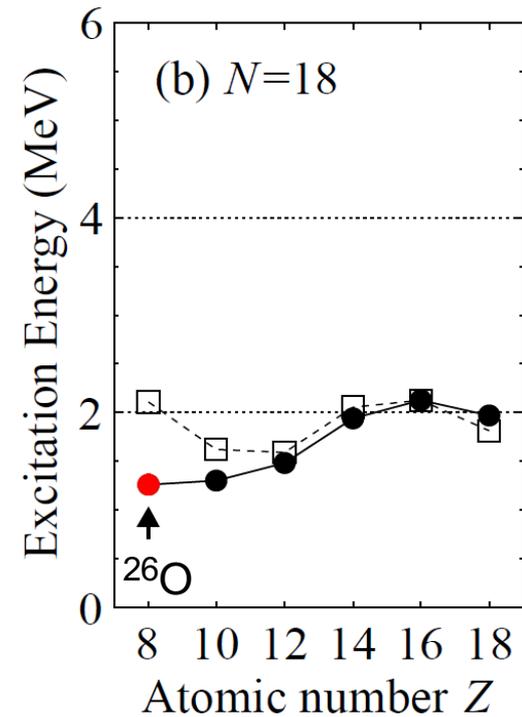
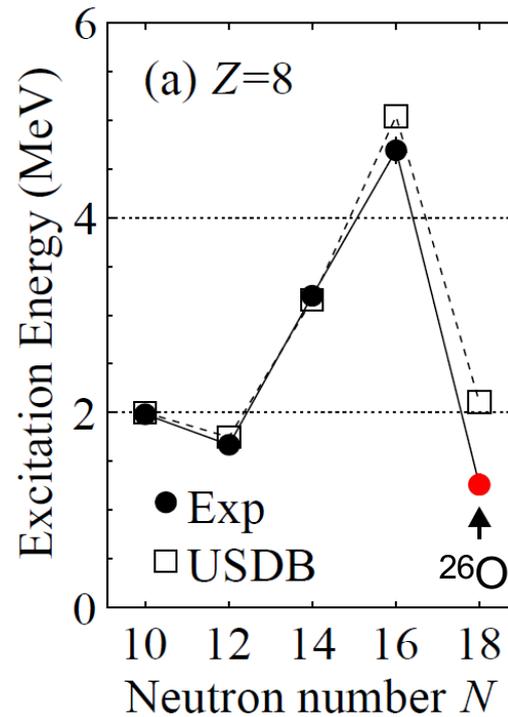
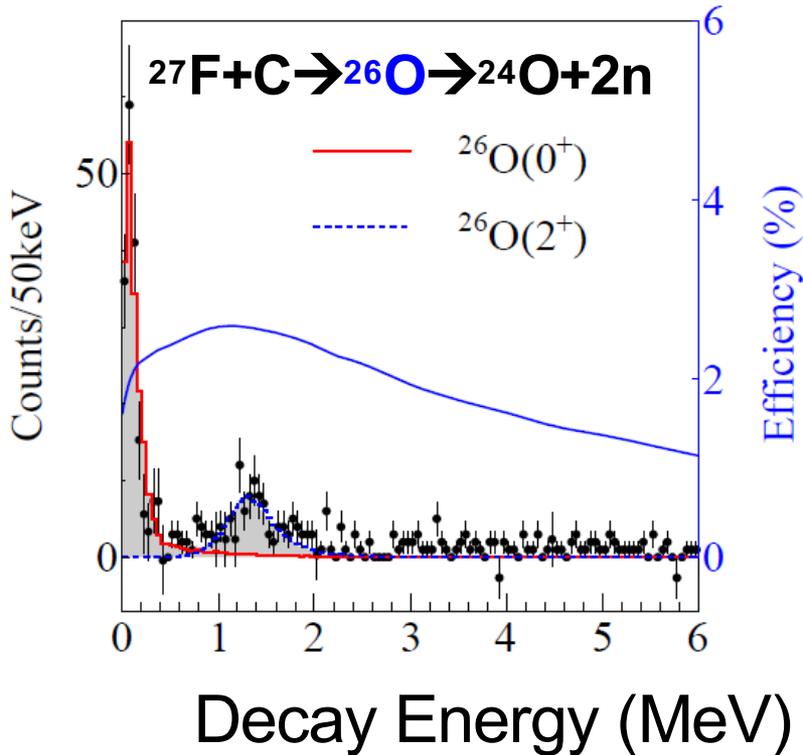
Large uncertainty of experimental study

- Only upper limit is given for the ground state energy
- Large systematic error in the lifetime measurement
- Excited State of ^{26}O ?

Experimental Setup at SAMURAI at RIBF



Study of ^{26}O (SAMURAI02)



Ground state (0^+)

5 times higher statistics than previous study

$18 \pm 3(\text{stat}) \pm 4(\text{syst})\text{keV}$

Finite value is determined for the first time

1st excited state (2^+)

Observed for the first time

$1.28^{+0.11}_{-0.08}\text{MeV}$

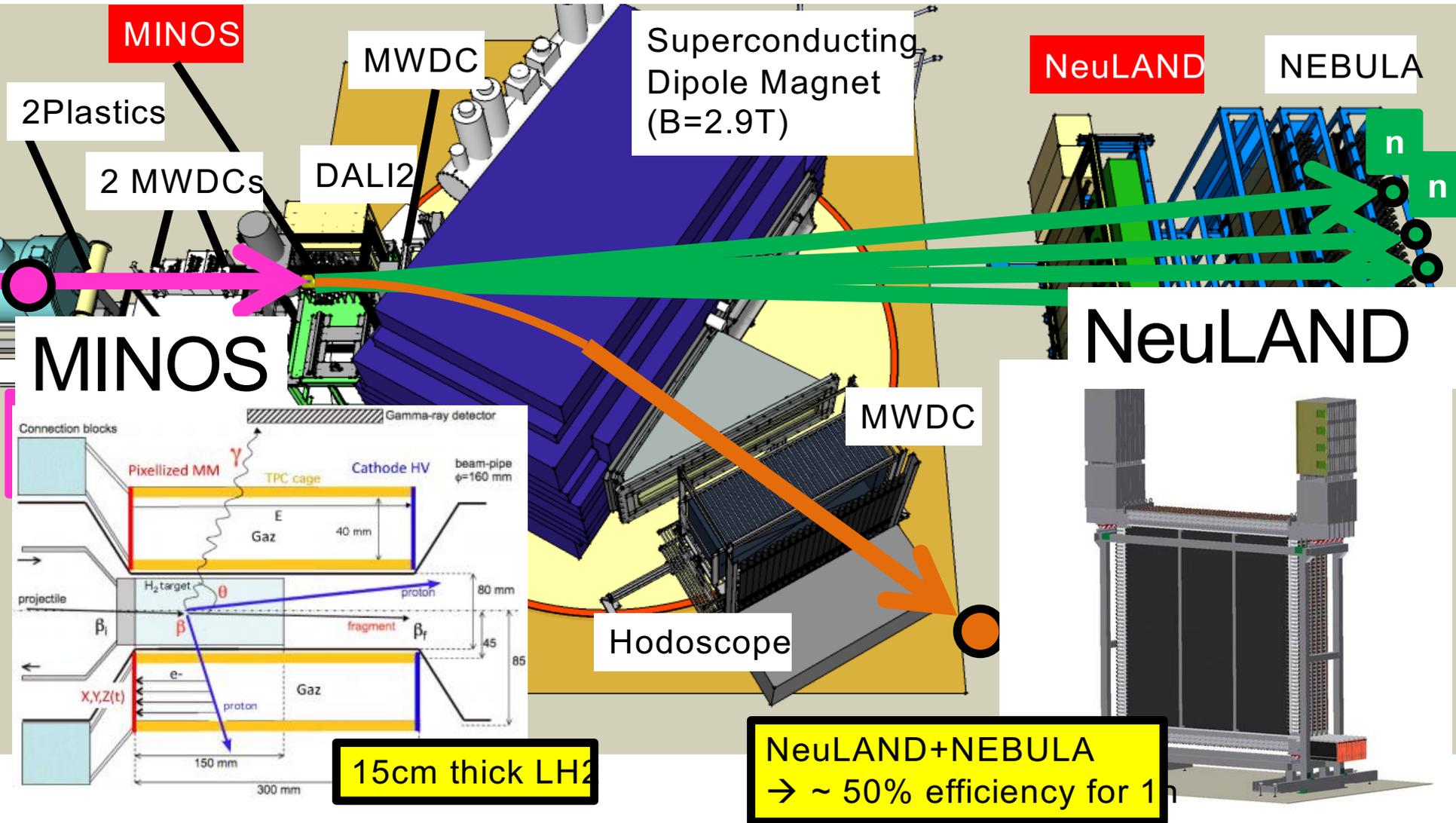
$N=16$ shell closure is confirmed

USDB cannot describe 2^+ energy at ^{26}O

→ effects of
pf shell?, continuum?

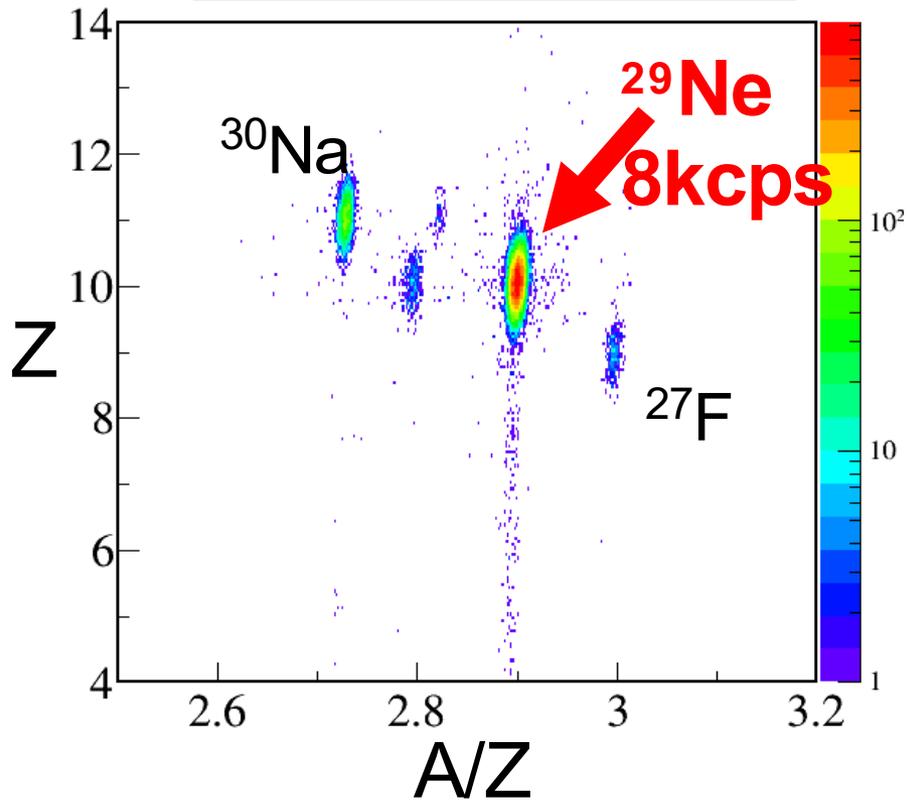
$2n$ Correlations?, $3N$ force?

^{28}O measurement @ RIBF-SAMURAI

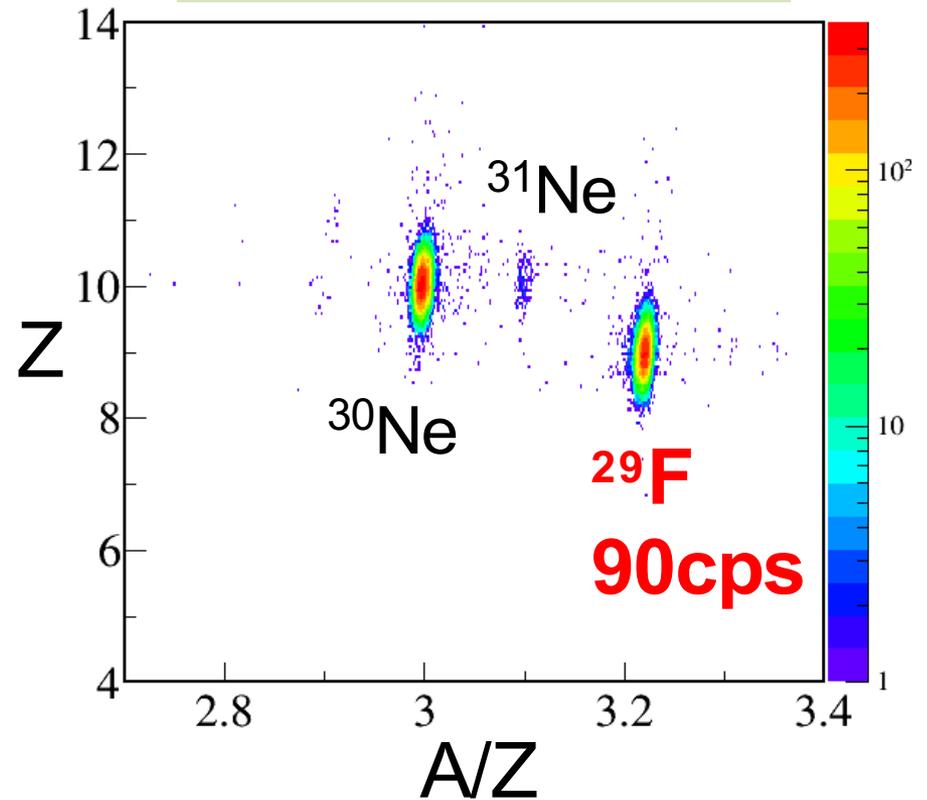


Secondary beam

^{29}Ne beam setting
(^{27}O measurement)



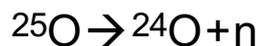
^{29}F beam setting
(^{28}O measurement)



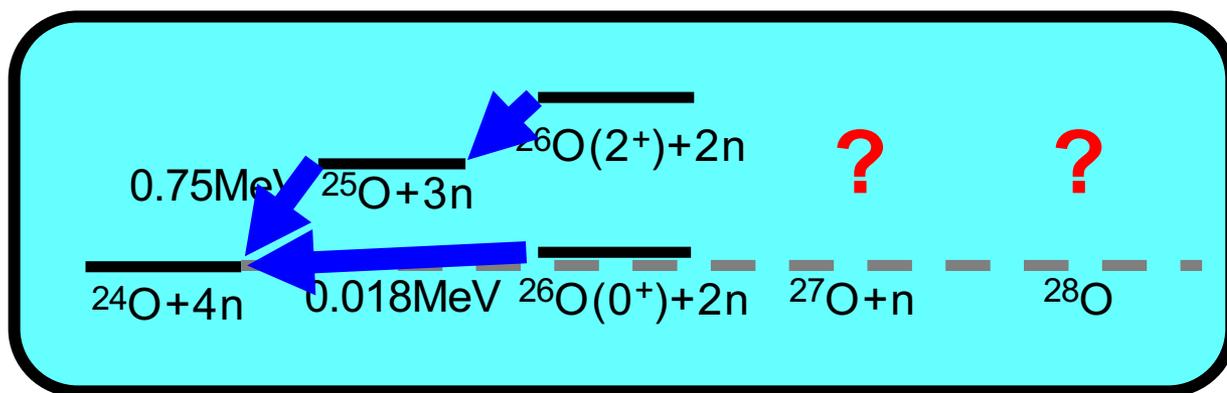
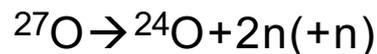
Preliminary decay energy spectra (subsystems, 1n/2n coincidence)

^{26}Ne	^{27}Ne	^{28}Ne	^{29}Ne	^{30}Ne	^{31}Ne
^{25}F	^{26}F	^{27}F	^{28}F	^{29}F	^{30}F
^{24}O	^{25}O	^{26}O	^{27}O	^{28}O	

Includes

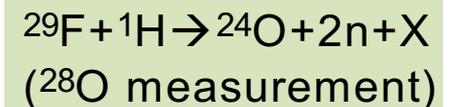
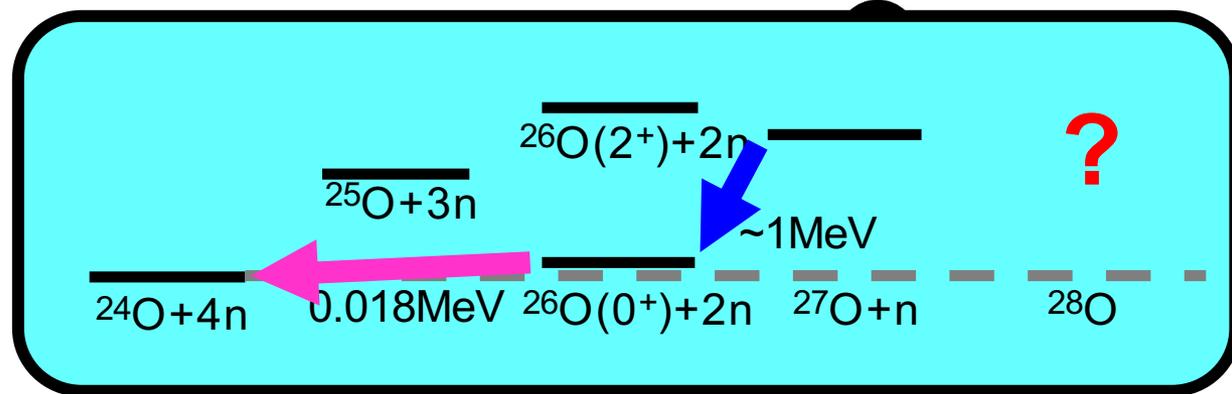
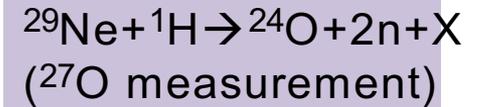


Includes



^{26}Ne	^{27}Ne	^{28}Ne	^{29}Ne	^{30}Ne	^{31}Ne
^{25}F	^{26}F	^{27}F	^{28}F	^{29}F	^{30}F
^{24}O	^{25}O	^{26}O	^{27}O	^{28}O	

Analysis of 2n coincidence events



**Will be checked from
analysis of 3n/4n
coincidence events**
Slide by Y.Kondo

● Summary and Outlook

□ SAMURAI at RIBF since 2012

- ✓ Large Acceptance, Multi-Purpose Spectrometer
with Neutron/Heavy ion/proton detectors
- ✓ NeuLAND(2015-2017) and MINOS(2014-2017)
→ Great Contributions to explore the extreme of nuclei
- ✓ A Variety of Opportunities for RI-Beam Science

□ Probes for extremes: Breakup, Quasi-free Scattering

□ Coulomb Breakup of Island-of-inversion nucleus: ^{31}Ne

□ Spectroscopy of $^{26-28}\text{O}$: Exploration beyond the drip line

^{26}O : Y. Kondo et al., PRL 116, 102503, (2016) .

Combination of the State-of-the-art-detectors
with Advanced Facilities

→ Essential for Exploring the Extremes of Nuclei

Near Future: → More Extremes are to be Explored !

Day-one Collaboration

Tokyo Institute of Technology: Y.Kondo, T.Nakamura, N.Kobayashi, R.Tanaka, R.Minakata, S.Ogoshi, S.Nishi, D.Kanno, T.Nakashima, J. Tsubota, A. Saito

LPC CAEN: N.A.Orr, J.Gibelin, F.Delaunay, F.M.Marques, N.L.Achouri, S.Lebond, O. Deshayes

Tohoku University : T.Koabayshi, K.Takahashi, K.Muto

RIKEN: K.Yoneda, T.Motobayashi ,H.Otsu, T.Isobe, H.Baba,H.Sato, Y.Shimizu, J.Lee, P.Doornenbal, S.Takeuchi, N.Inabe, N.Fukuda, D.Kameda, H.Suzuki, H.Takeda, T.Kubo

Seoul National University: Y.Satou, S.Kim, J.W.Hwang

Kyoto University : T.Murakami, N.Nakatsuka

GSI : Y.Togano

Univ. of York: A.G.Tuff

GANIL: A.Navin

Technische Universit  at Darmstadt: T.Aumann

Rikkyo Univeristy: D.Murai

Universit  e Paris-Sud, IN2P3-CNRS: M.Vandebrouck

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